


OPTIMAL ERROR ESTIMATES OF A CRANK–NICOLSON FINITE ELEMENT PROJECTION METHOD FOR MAGNETOHYDRODYNAMIC EQUATIONS

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Abstract. In this paper, we propose and analyze a fully discrete finite element projection method for the magnetohydrodynamic (MHD) equations. A modified Crank–Nicolson method and the Galerkin finite element method are used to discretize the model in time and space, respectively, and appropriate semi-implicit treatments are applied to the fluid convection term and two coupling terms. These semi-implicit approximations result in a linear system with variable coefficients for which the unique solvability can be proved theoretically. In addition, we use a second-order decoupling projection method of the Van Kan type [Van Kan, *SIAM J. Sci. Statist. Comput.* **7** (1986) 870–891] in the Stokes solver, which computes the intermediate velocity field based on the gradient of the pressure from the previous time level, and enforces the incompressibility constraint *via* the Helmholtz decomposition of the intermediate velocity field. The energy stability of the scheme is theoretically proved, in which the decoupled Stokes solver needs to be analyzed in details. Error estimates are proved in the discrete $L^\infty(0, T; L^2)$ norm for the proposed decoupled finite element projection scheme. Numerical examples are provided to illustrate the theoretical results.

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1. INTRODUCTION

The magnetohydrodynamic equations have been widely applied into metallurgy and liquid-metal processing, and the numerical solutions are of great significance in practical scientific and engineering applications; see Asai [2] and Unger *et al.* [38]. Such an MHD system could be formulated as [35]

$$\mu \partial_t \mathbf{H} + \sigma^{-1} \nabla \times (\nabla \times \mathbf{H}) - \mu \nabla \times (\mathbf{u} \times \mathbf{H}) = \sigma^{-1} \nabla \times \mathbf{J}, \quad (1.1)$$

$$\partial_t \mathbf{u} + \mathbf{u} \cdot \nabla \mathbf{u} - \nu \Delta \mathbf{u} + \nabla p = \mathbf{f} - \mu \mathbf{H} \times (\nabla \times \mathbf{H}), \quad (1.2)$$

$$\nabla \cdot \mathbf{u} = 0, \quad (1.3)$$

over $\Omega \times (0, T]$, where Ω is a bounded and convex polyhedral domain in \mathbb{R}^3 (polygonal domain in \mathbb{R}^2). In the above system, \mathbf{u} , \mathbf{H} and p denote the velocity field, the magnetic field, and the pressure, respectively; \mathbf{J} and

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\mathbf{f} are the given source terms (\mathbf{J} denotes a scalar function in \mathbb{R}^2); σ denotes the magnetic Reynolds number, ν denotes the viscosity of the fluid, and $\mu = M^2\nu\sigma^{-1}$, where M is the Hartman number. The initial data and boundary conditions are given by

$$\mathbf{H}|_{t=0} = \mathbf{H}_0, \quad \mathbf{u}|_{t=0} = \mathbf{u}_0, \quad \text{in } \Omega, \quad (1.4)$$

$$\mathbf{H} \times \mathbf{n} = \mathbf{0}, \quad \mathbf{u} = \mathbf{0}, \quad \text{on } \partial\Omega \times (0, T]. \quad (1.5)$$

It is assumed that the initial data satisfies

$$\nabla \cdot \mathbf{H}_0 = \nabla \cdot \mathbf{u}_0 = 0. \quad (1.6)$$

By taking the divergence of (1.1), one can easily get $\mu\partial_t \nabla \cdot \mathbf{H} = 0$, which together with the above divergence-free initial condition implies that

$$\nabla \cdot \mathbf{H} = 0. \quad (1.7)$$

The existence and uniqueness of the weak solution for this problem has been theoretically proved in [15, 33]. More regularity analysis of the MHD system could be referred in [17, 25, 26, 32], etc.

There have been many existing works on the numerical approximations for the incompressible MHD system. In bounded and convex domains, the solutions of the MHD model are generally in $[H^1(\Omega)]^d$ ($d = 2, 3$, denotes the dimension of Ω) and therefore people often use H^1 -conforming finite element methods (FEMs) to solve the MHD equations numerically. For example, Gunzburger *et al.* [15] used H^1 -conforming FEMs for solving the stationary incompressible MHD equations with an optimal error estimate being established. Later, He developed H^1 -conforming FEMs in [18] for solving the time-dependent MHD equations and proved error estimates of the numerical scheme. More works on H^1 -conforming FEMs for the MHD equations can be found in [1, 10, 13, 19, 24].

While the spatial approximation has always been important, the temporal discretization also plays a significant role for solving the MHD system. There have been quite a few existing stability and convergence analyses for the first-order temporally accurate numerical schemes [9, 18, 28–30]. In most of these works, the stability and convergence analyses have been based on a Stokes solver at each time step, *i.e.*, the computation of the pressure gradient has to be implemented with the incompressibility constraint being enforced, which in turn leads to a non-symmetric linear system, and the computation costs turn out to be extremely expensive. To overcome this difficulty, some “decoupled” techniques have been introduced. In [44], Zhao *et al.* dealt with a binary hydrodynamic phase field model of mixtures of nematic liquid crystals and viscous fluids by designing a decoupled semi-discrete scheme, which is linear, first-order accurate in time, and unconditionally energy stable. In particular, a pressure-correction scheme [14] was used so that the pressure could be explicitly updated in the velocity equation by introducing an intermediate function and thus two sub-systems are generated. In [27], Liu *et al.* proposed a first-order decoupled scheme for a phase-field model of two-phase incompressible flows with variable density based on a “pressure-stabilized” formulation, which treats the pressure term explicitly in the velocity field equation, and only requires a Poisson solver to update the pressure. These works have mainly focused on the design of energy-preserving schemes without presenting the convergence analysis. Meanwhile, the first-order temporal accuracy may not be sufficient in the practical computations of the MHD system, and therefore higher-order temporal numerical approximations have been highly desired.

In the development of temporally higher-order methods, a conditionally stable second-order backward difference formula (BDF2) algorithm was proposed in [23] for a reduced MHD model at small magnetic Reynolds number, in which the coupling terms were explicitly updated, and other terms were implicitly computed. An unconditionally stable BDF2 method was proposed in [20], where the method was proved convergent with optimal order. In [43], a second-order scheme with Newton treatment of the nonlinear terms was proposed, where the unconditional stability and optimal error estimates were obtained. Recently, a fully discrete Crank–Nicolson (CN) scheme was studied in [21], where the unconditional energy stability and convergence (without error estimates) were proved. For efficient large scale numerical simulations of incompressible flows, high-order projection

methods are desired. In [34], Shen presented rigorous error analysis of second-order Crank–Nicolson projection methods of the Van Kan type [39], *i.e.*, second-order incremental pressure-correction methods, for the unsteady incompressible Navier–Stokes equations. By interpreting the respective projections schemes as second-order time discretizations of a perturbed system which approximates the Navier–Stokes equations, optimal-order convergence in the discrete $L^2(0, T; L^2)$ norm was proved for the semi-implicit schemes. Later, Guermond [12] proved optimal error estimates in the discrete $L^2(0, T; L^2)$ norm for the fully discrete case with BDF2 approximation in time. However, whether second-order incremental pressure-correction methods have optimal convergence in the discrete $L^\infty(0, T; L^2)$ norm remains open for both Navier–Stokes and MHD equations.

In this article, we fill in the gap between numerical computation and rigorous error estimates for a Crank–Nicolson finite element projection method for the MHD model. We first propose a fully discrete decoupled finite element projection method for the MHD system (1.1)–(1.3), and then the following properties are theoretically established: unique solvability, unconditional energy stability, and convergence analysis. In particular, a modified Crank–Nicolson method with an implicit Adams–Moulton interpolation in the form of $\frac{3}{4}\mathbf{H}_h^{n+1} + \frac{1}{4}\mathbf{H}_h^{n-1}$, instead of the standard Crank–Nicolson approximation, is applied to discretize the magnetic diffusion term. Such a technique leads to a stronger stability property of the numerical scheme, as will be demonstrated in the subsequent analysis.

A second-order incremental pressure-correction method is used to decouple the computation of velocity and pressure. Precisely, an intermediate velocity function $\widehat{\mathbf{u}}_h^{n+1}$ is introduced in the numerical scheme, and its computation is based on the pressure gradient at the previous time step. After solving the intermediate velocity field, we decompose it into the divergence-free subspace by using the Helmholtz decomposition. This yields the velocity field \mathbf{u}_h^{n+1} at the same time level. In the error analysis, we first introduce an intermediate projection of the velocity, *i.e.*, $\widehat{\mathbf{R}}_h \widehat{\mathbf{u}}_h^{n+1}$ as introduced in (4.16) and (4.17), with which, estimate of an intermediate error for the velocity is obtained. With such estimate and rigorous analysis of the discrete gradient of the Stokes projection, second-order convergence in time is proved in the discrete $L^\infty(0, T; L^2)$ norm for the velocity and magnetic fields, independently of the mesh size h in the case $\nabla p|_{\partial\Omega} = 0$ and dependently on $h^{-\frac{1}{2}}$ in the case $\nabla p|_{\partial\Omega} \neq 0$, respectively. The techniques introduced in this paper would also work for other related projection methods.

This paper is organized as follows. In Section 2, a variational formulation and some preliminary results are reviewed. The fully discrete finite element scheme is introduced in Section 3, and its unconditional energy stability is established in details. Section 4 provides the rigorous proof of the unique solvability and error estimates. Several numerical examples are presented in Section 5. Finally, some concluding remarks are provided in Section 6.

2. VARIATIONAL FORMULATION AND STABILITY ANALYSIS

For $k \geq 0$ and $1 \leq p \leq \infty$, let $W^{k,p}(\Omega)$ be the conventional Sobolev space of functions defined on Ω , with abbreviations $L^p(\Omega) = W^{0,p}(\Omega)$ and $H^k(\Omega) = W^{k,2}(\Omega)$. Then, we denote by $W_0^{1,p}(\Omega)$ the space of functions in $W^{1,p}(\Omega)$ with zero traces on the boundary $\partial\Omega$, and denote $H_0^1(\Omega) = W_0^{1,2}(\Omega)$. The corresponding vector-valued spaces are

$$\begin{aligned} \mathbf{L}^p(\Omega) &= [L^p(\Omega)]^d, & \mathbf{W}^{k,p}(\Omega) &= [W^{k,p}(\Omega)]^d, \\ \mathbf{W}_0^{1,p}(\Omega) &= [W_0^{1,p}(\Omega)]^d, & \mathbf{H}_0^1(\Omega) &= \mathbf{W}_0^{1,2}(\Omega), \\ \mathring{\mathbf{H}}^1(\Omega) &= \{\mathbf{v} \in \mathbf{H}^1(\Omega) : \mathbf{v} \times \mathbf{n}|_{\partial\Omega} = 0\}, \end{aligned}$$

where $d = 2, 3$, denotes the dimension of Ω . As usual, the inner product of $L^2(\Omega)$ is denoted by (\cdot, \cdot) .

With the above notations, it could be seen that the exact solution $(\mathbf{H}, \mathbf{u}, p)$ of (1.1)–(1.3) satisfies

$$(\mu \partial_t \mathbf{H}, \mathbf{w}) + (\sigma^{-1} \nabla \times \mathbf{H}, \nabla \times \mathbf{w}) - (\mu \mathbf{u} \times \mathbf{H}, \nabla \times \mathbf{w}) = (\sigma^{-1} \nabla \times \mathbf{J}, \mathbf{w}), \quad (2.1)$$

$$(\partial_t \mathbf{u}, \mathbf{v}) + (\nu \nabla \mathbf{u}, \nabla \mathbf{v}) + b(\mathbf{u}, \mathbf{u}, \mathbf{v}) - (p, \nabla \cdot \mathbf{v}) = (\mathbf{f}, \mathbf{v}) - (\mu \mathbf{H} \times (\nabla \times \mathbf{H}), \mathbf{v}), \quad (2.2)$$

$$(\nabla \cdot \mathbf{u}, q) = 0, \quad (2.3)$$

for any test functions $(\mathbf{w}, \mathbf{v}, q) \in (\mathring{\mathbf{H}}^1(\Omega), \mathbf{H}_0^1(\Omega), L^2(\Omega))$, where we have defined the trilinear form $b(\cdot, \cdot, \cdot)$ as

$$\begin{aligned} b(\mathbf{u}, \mathbf{v}, \mathbf{w}) &:= (\mathbf{u} \cdot \nabla \mathbf{v}, \mathbf{w}) + \frac{1}{2}((\nabla \cdot \mathbf{u})\mathbf{v}, \mathbf{w}) \\ &= \frac{1}{2}[(\mathbf{u} \cdot \nabla \mathbf{v}, \mathbf{w}) - (\mathbf{u} \cdot \nabla \mathbf{w}, \mathbf{v})], \quad \forall \mathbf{u}, \mathbf{v}, \mathbf{w} \in \mathbf{H}_0^1(\Omega), \end{aligned} \quad (2.4)$$

and $\mathbf{u} \cdot \mathbf{v}$ denotes the Euclidean scalar product in \mathbb{R}^d . Notice that the trilinear form $b(\cdot, \cdot, \cdot)$ is skew-symmetric with respect to its last two arguments, so that we further have

$$b(\mathbf{u}, \mathbf{v}, \mathbf{v}) = 0, \quad \forall \mathbf{u}, \mathbf{v}, \mathbf{w} \in \mathbf{H}_0^1(\Omega).$$

The energy stability of the continuous system (2.1)–(2.3) could be obtained in a straightforward manner. By taking $\mathbf{w} = \mathbf{H}$, $\mathbf{v} = \mathbf{u}$ in (2.1)–(2.3) and adding the resulting equations together, we get

$$\begin{aligned} \frac{\mu}{2} \frac{d}{dt} \|\mathbf{H}\|_{L^2}^2 + \sigma^{-1} \|\nabla \times \mathbf{H}\|_{L^2}^2 + \frac{1}{2} \frac{d}{dt} \|\mathbf{u}\|_{L^2}^2 + \nu \|\nabla \mathbf{u}\|_{L^2}^2 &= (\mathbf{J}, \sigma^{-1} \nabla \times \mathbf{H}) + (\mathbf{f}, \mathbf{u}) \\ &\leq \frac{1}{4\sigma} \|\mathbf{J}\|_{L^2}^2 + \sigma^{-1} \|\nabla \times \mathbf{H}\|_{L^2}^2 + \frac{1}{4\varepsilon} \|\mathbf{f}\|_{L^2}^2 + \varepsilon \|\mathbf{u}\|_{L^2}^2, \end{aligned}$$

where ε is an arbitrary constant. Due to the zero boundary condition of \mathbf{u} in (1.5), we have $\|\mathbf{u}\|_{L^2}^2 \leq C \|\nabla \mathbf{u}\|_{L^2}^2$. Since ε can be arbitrarily small, we obtain the following energy estimate

$$\frac{\mu}{2} \frac{d}{dt} \|\mathbf{H}\|_{L^2}^2 + \frac{1}{2} \frac{d}{dt} \|\mathbf{u}\|_{L^2}^2 \leq \frac{1}{4\sigma} \|\mathbf{J}\|_{L^2}^2 + \frac{1}{4\varepsilon} \|\mathbf{f}\|_{L^2}^2. \quad (2.5)$$

If the sources terms $\mathbf{J} = \mathbf{f} = \mathbf{0}$, we further get

$$\frac{\mu}{2} \frac{d}{dt} \|\mathbf{H}\|_{L^2}^2 + \frac{1}{2} \frac{d}{dt} \|\mathbf{u}\|_{L^2}^2 \leq 0, \quad (2.6)$$

which implies the total energy is decaying.

3. NUMERICAL METHOD AND THEORETICAL RESULTS

3.1. Numerical method

In this subsection, we propose a fully discrete decoupled finite element method for solving the system (1.1)–(1.3). Let \mathfrak{S}_h denote a quasi-uniform partition of Ω into tetrahedrons K_j in \mathbb{R}^3 (or triangles in \mathbb{R}^2), $j = 1, 2, \dots, M$, with mesh size $h = \max_{1 \leq j \leq M} \{\text{diam} K_j\}$. To approximate \mathbf{u} and p in the system (1.1)–(1.3), we introduce the Taylor-Hood finite element space $\mathbf{X}_h \times M_h$, defined by

$$\begin{aligned} \mathbf{X}_h &= \{\mathbf{l}_h \in \mathbf{H}_0^1(\Omega) : \mathbf{l}_h|_{K_j} \in \mathbf{P}_r(K_j)\}, \\ M_h &= \left\{ q_h \in L^2(\Omega) : q_h|_{K_j} \in P_{r-1}(K_j), \int_{\Omega} q_h dx = 0 \right\}, \end{aligned}$$

for any integer $r \geq 2$, where $P_r(K_j)$ is the space of polynomials with degree r on K_j for all $K_j \in \mathfrak{S}_h$ and $\mathbf{P}_r(K_j) := [P_r(K_j)]^d$. To approximate the magnetic field \mathbf{H} , we introduce the finite element space \mathbf{S}_h defined by

$$\mathbf{S}_h = \{\mathbf{w}_h \in \mathring{\mathbf{H}}^1(\Omega) : \mathbf{w}_h|_{K_j} \in \mathbf{P}_r(K_j)\}.$$

Let $\{t_n = n\tau\}_{n=0}^N$ denote a uniform partition of the time interval $[0, T]$, with a step size $\tau = T/N$, and $v^n = v(x, t_n)$. For any sequences $\{v^n\}_{n=0}^N$ and $\{\hat{v}^n\}_{n=0}^N$, we define

$$\tilde{v}^{n+\frac{1}{2}} := \frac{3}{4}v^{n+1} + \frac{1}{4}v^{n-1}, \quad \bar{v}^{n+\frac{1}{2}} := \frac{1}{2}\hat{v}^{n+1} + \frac{1}{2}v^n, \quad \tilde{v}^{n+\frac{1}{2}} := \frac{3}{2}v^n - \frac{1}{2}v^{n-1}.$$

Then, a fully discrete decoupled Crank–Nicolson finite element projection method for the incompressible MHD equations (1.1)–(1.3) is formulated as: find $(\mathbf{H}_h^{n+1}, \mathbf{u}_h^{n+1}, \hat{\mathbf{u}}_h^{n+1}, p_h^{n+1}) \in (\mathbf{S}_h, \mathbf{X}_h, \mathbf{X}_h, M_h)$ such that

$$\begin{aligned} \mu \left(\frac{\mathbf{H}_h^{n+1} - \mathbf{H}_h^n}{\tau}, \mathbf{w}_h \right) + \sigma^{-1} \left(\nabla \times \tilde{\mathbf{H}}_h^{n+\frac{1}{2}}, \nabla \times \mathbf{w}_h \right) + \sigma^{-1} \left(\nabla \cdot \tilde{\mathbf{H}}_h^{n+\frac{1}{2}}, \nabla \cdot \mathbf{w}_h \right) \\ - \mu \left(\bar{\mathbf{u}}_h^{n+\frac{1}{2}} \times \tilde{\mathbf{H}}_h^{n+\frac{1}{2}}, \nabla \times \mathbf{w}_h \right) = \sigma^{-1} \left(\nabla \times \mathbf{J}^{n+\frac{1}{2}}, \mathbf{w}_h \right), \end{aligned} \quad (3.1)$$

$$\begin{aligned} \left(\frac{\hat{\mathbf{u}}_h^{n+1} - \mathbf{u}_h^n}{\tau}, \mathbf{v}_h \right) + \nu \left(\nabla \bar{\mathbf{u}}_h^{n+\frac{1}{2}}, \nabla \mathbf{v}_h \right) + b \left(\tilde{\mathbf{u}}_h^{n+\frac{1}{2}}, \bar{\mathbf{u}}_h^{n+\frac{1}{2}}, \mathbf{v}_h \right) - (p_h^n, \nabla \cdot \mathbf{v}_h) \\ + \mu \left(\tilde{\mathbf{H}}_h^{n+\frac{1}{2}} \times \left(\nabla \times \tilde{\mathbf{H}}_h^{n+\frac{1}{2}} \right), \mathbf{v}_h \right) = \left(\mathbf{f}^{n+\frac{1}{2}}, \mathbf{v}_h \right), \end{aligned} \quad (3.2)$$

$$\left(\frac{\mathbf{u}_h^{n+1} - \hat{\mathbf{u}}_h^{n+1}}{\tau}, \mathbf{l}_h \right) - \frac{1}{2} (p_h^{n+1} - p_h^n, \nabla \cdot \mathbf{l}_h) = 0, \quad (3.3)$$

$$(\nabla \cdot \mathbf{u}_h^{n+1}, q_h) = 0, \quad (3.4)$$

for any $(\mathbf{w}_h, \mathbf{v}_h, \mathbf{l}_h, q_h) \in (\mathbf{S}_h, \mathbf{X}_h, \mathbf{X}_h, M_h)$ and $n = 1, 2, \dots, N-1$. Here we have added a stabilization term $\sigma^{-1}(\nabla \cdot \tilde{\mathbf{H}}_h^{n+\frac{1}{2}}, \nabla \cdot \mathbf{w}_h)$ to (3.1). This is consistent with (2.1) in view of (1.7).

In this paper, it is assumed that the system (1.1)–(1.3) admits a unique solution satisfying

$$\begin{aligned} \|\mathbf{H}_{ttt}\|_{L^\infty(0,T;L^2)} + \|\mathbf{H}_{tt}\|_{L^\infty(0,T;H^1)} + \|\mathbf{H}_t\|_{L^\infty(0,T;H^{r+1})} + \|\mathbf{u}_{ttt}\|_{L^\infty(0,T;L^2)} \\ + \|\mathbf{u}_{tt}\|_{L^\infty(0,T;H^1)} + \|\mathbf{u}_t\|_{L^\infty(0,T;H^{r+1})} + \|p_{tt}\|_{L^\infty(0,T;L^2)} + \|p_t\|_{L^\infty(0,T;H^r)} \leq K. \end{aligned} \quad (3.5)$$

Here, the subscripts of $\mathbf{H}, \mathbf{u}, p$ denote the partial derivative to variable t .

Next, we present our main results, *i.e.*, error estimates for scheme (3.1)–(3.4), in the following theorem.

Theorem 3.1. *Suppose that the system (1.1)–(1.3) has a unique solution $(\mathbf{H}, \mathbf{u}, p)$ satisfying (3.5). Then there exist positive constants τ_0 and h_0 such that when $\tau < \tau_0$, $h < h_0$, and $\tau = \mathcal{O}(h)$, the fully discrete decoupled FEM system (3.1)–(3.4) admits a unique solution $(\mathbf{H}_h^n, \mathbf{u}_h^n, p_h^n)$, $n = 2, 3, \dots, N$, which satisfies that*

$$\max_{2 \leq n \leq N} (\|\mathbf{H}_h^n - \mathbf{H}^n\|_{L^2} + \|\mathbf{u}_h^n - \mathbf{u}^n\|_{L^2}) \leq C_0 (\ell_h \tau^2 + h^{r+1}), \quad (3.6)$$

$$\left(\tau \sum_{n=2}^N \left(\|\nabla \times (\mathbf{H}_h^n - \mathbf{H}^n)\|_{L^2}^2 + \|\nabla \left(\bar{\mathbf{u}}_h^{n-\frac{1}{2}} - \bar{\mathbf{u}}^{n-\frac{1}{2}} \right)\|_{L^2}^2 \right) \right)^{\frac{1}{2}} \leq C_0 (\ell_h \tau^2 + h^r), \quad (3.7)$$

where $\bar{\mathbf{u}}^{n-\frac{1}{2}} := \frac{1}{2}(\hat{\mathbf{u}}^n + \mathbf{u}^{n-1})$ and $\hat{\mathbf{u}}^n := \mathbf{u}^n$; $\ell_h = 1$ if $\nabla p|_{\partial\Omega} = \mathbf{0}$, otherwise $\ell_h = h^{-\frac{1}{2}}$; C_0 is a positive constant independent of τ and h .

Remark 3.2. One feature of the proposed numerical scheme (3.1)–(3.4) is associated with its decoupled nature in the Stokes solver. Motivated by the second-order projection method of the Van Kan type [39], *i.e.*, second-order incremental pressure-correction method, we introduce an intermediate velocity $\hat{\mathbf{u}}_h^{n+1}$ to decouple the problem, and thus build two systems and both of them consist of two unknowns. More precisely, we first obtain \mathbf{H}_h^{n+1} and $\hat{\mathbf{u}}_h^{n+1}$ through (3.1) and (3.2), while treating the gradient of pressure explicitly. Then, we substitute

$\hat{\mathbf{u}}_h^{n+1}$ into (3.3) and (3.4), so that p_h^{n+1} and \mathbf{u}_h^{n+1} could be efficiently computed *via* solving a Darcy problem. In comparison with the classical coupled solver that the full system contains three unknowns \mathbf{H}_h^{n+1} , \mathbf{u}_h^{n+1} and p_h^{n+1} , which have to be solved simultaneously, such a decoupled approach will greatly improve the efficiency of the numerical scheme.

There have been extensive analyses of decoupled numerical schemes for incompressible Navier–Stokes equations; see the pioneering works of Chorin [7], Temam [36], and many other related studies [3, 22, 31, 40–42], etc. In this work, second-order convergence in time is proved in the discrete $L^\infty(0, T; L^2)$ norm for the velocity and magnetic fields, independently of the mesh size h in the case $\nabla p|_{\partial\Omega} = 0$ and dependently on $h^{-\frac{1}{2}}$ in the case $\nabla p|_{\partial\Omega} \neq 0$, respectively. The techniques introduced in this paper would also work for other related projection methods.

Remark 3.3. Another feature of scheme (3.1)–(3.4) is that we have used a modified Crank–Nicolson method for temporal discretization, where the term $\nabla \times \mathbf{H}^{n+\frac{1}{2}}$ is approximated by $\nabla \times (\frac{3}{4}\mathbf{H}^{n+1} + \frac{1}{4}\mathbf{H}^{n-1})$. This enables us to obtain error estimates for the term $\nabla \times \mathbf{H}$ at certain time steps, instead of an average of those at two consecutive time levels; see (3.7). Such a modified Crank–Nicolson scheme has been extensively applied to various gradient flow models [5, 6, 8, 16]. An application of this approach to the incompressible MHD system is reported in this work, for the first time.

Remark 3.4. In (3.1), we have added a stabilization term $\sigma^{-1}(\nabla \cdot \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}}, \nabla \cdot \mathbf{w}_h)$ to validate the coercivity of the magnetic equation, with which, optimal error estimates for the magnetic field in energy-norm can be proved.

Remark 3.5. It is noted that the numerical solutions at two previous time levels are needed for the implementation of (3.1)–(3.4). The starting values at time steps t_0 and t_1 are assumed to be given and satisfy the estimate (3.6). An example of constructing the numerical schemes for starting values is an application of the backward Euler FEM method at t_1 and the Stokes and Maxwell projections of the initial data at t_0 . In such case, the error estimate (3.6) holds at t_1 and t_0 .

In the following subsection, we analyze the energy stability of scheme (3.1)–(3.4). In this paper, we denote by C a generic positive constant and by ε a generic small positive constant, which are independent of n , h , τ , and C_0 .

3.2. Stability analysis of numerical scheme

In this subsection, we present the energy stability analysis for the numerical system (3.1)–(3.4). Here, we introduce a discrete version of the gradient operator, $\nabla_h : M_h \rightarrow \mathbf{X}_h$, defined as

$$(\mathbf{v}_h, \nabla_h q_h) = -(\nabla \cdot \mathbf{v}_h, q_h), \quad \forall \mathbf{v}_h \in \mathbf{X}_h, q_h \in M_h. \quad (3.8)$$

Through the definition of the discrete gradient operator ∇_h , we can rewrite the equation (3.3) in the following equivalent form:

$$\frac{\mathbf{u}_h^{n+1} - \hat{\mathbf{u}}_h^{n+1}}{\tau} + \frac{1}{2}\nabla_h(p_h^{n+1} - p_h^n) = 0. \quad (3.9)$$

The abstract form (3.9) will be useful in the stability analysis of numerical scheme.

Theorem 3.6. *The numerical solution $(\mathbf{H}_h^n, \mathbf{u}_h^n, p_h^n)$ to the fully discrete linearized FEM (3.1)–(3.4) satisfies the following energy stability estimate*

$$\frac{\mu}{2\tau} \left(\|\mathbf{H}_h^{n+1}\|_{L^2}^2 - \|\mathbf{H}_h^n\|_{L^2}^2 \right) + \frac{\mu}{8\tau} \left(\|\mathbf{H}_h^{n+1} - \mathbf{H}_h^n\|_{L^2}^2 - \|\mathbf{H}_h^n - \mathbf{H}_h^{n-1}\|_{L^2}^2 \right)$$

$$\begin{aligned}
& + \frac{1}{2\tau} \left(\|\mathbf{u}_h^{n+1}\|_{L^2}^2 - \|\mathbf{u}_h^n\|_{L^2}^2 \right) + \frac{\tau}{8} \left(\|\nabla_h p_h^{n+1}\|_{L^2}^2 - \|\nabla_h p_h^n\|_{L^2}^2 \right) \\
& \leq C \left(\|\mathbf{J}^{n+\frac{1}{2}}\|_{L^2}^2 + \|\mathbf{f}^{n+\frac{1}{2}}\|_{L^2}^2 \right),
\end{aligned} \tag{3.10}$$

for $n = 1, 2, \dots, N-1$, where C is a positive constant independent of τ and h .

Proof. By taking $\mathbf{w}_h = \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}}$ in (3.1) and $\mathbf{v}_h = \overline{\mathbf{u}}_h^{n+\frac{1}{2}}$ in (3.2), we get

$$\begin{aligned}
& \frac{\mu}{2\tau} \left(\|\mathbf{H}_h^{n+1}\|_{L^2}^2 - \|\mathbf{H}_h^n\|_{L^2}^2 \right) + \frac{\mu}{8\tau} \left(\|\mathbf{H}_h^{n+1} - \mathbf{H}_h^n\|_{L^2}^2 - \|\mathbf{H}_h^n - \mathbf{H}_h^{n-1}\|_{L^2}^2 \right) \\
& + \frac{\mu}{8\tau} \|\mathbf{H}_h^{n+1} - 2\mathbf{H}_h^n + \mathbf{H}_h^{n-1}\|_{L^2}^2 + \sigma^{-1} \|\nabla \times \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}}\|_{L^2}^2 + \sigma^{-1} \|\nabla \cdot \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}}\|_{L^2}^2 \\
& - \mu \left(\overline{\mathbf{u}}_h^{n+\frac{1}{2}} \times \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}}, \nabla \times \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \right) \\
& = \sigma^{-1} \left(\nabla \times \mathbf{J}^{n+\frac{1}{2}}, \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \right),
\end{aligned} \tag{3.11}$$

and

$$\begin{aligned}
& \frac{1}{2\tau} \left(\|\widehat{\mathbf{u}}_h^{n+1}\|_{L^2}^2 - \|\mathbf{u}_h^n\|_{L^2}^2 \right) + \nu \|\nabla \overline{\mathbf{u}}_h^{n+\frac{1}{2}}\|_{L^2}^2 - \left(p_h^n, \nabla \cdot \overline{\mathbf{u}}_h^{n+\frac{1}{2}} \right) + \mu \left(\widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \times \left(\nabla \times \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \right), \overline{\mathbf{u}}_h^{n+\frac{1}{2}} \right) \\
& = \left(\mathbf{f}^{n+\frac{1}{2}}, \overline{\mathbf{u}}_h^{n+\frac{1}{2}} \right),
\end{aligned} \tag{3.12}$$

respectively, where we have used the fact that $b(\widehat{\mathbf{u}}_h^{n+\frac{1}{2}}, \overline{\mathbf{u}}_h^{n+\frac{1}{2}}, \overline{\mathbf{u}}_h^{n+\frac{1}{2}}) = 0$, and

$$\begin{aligned}
\left(\frac{\mathbf{H}_h^{n+1} - \mathbf{H}_h^n}{\tau}, \mathbf{w}_h \right) & = \frac{1}{2\tau} \left(\|\mathbf{H}_h^{n+1}\|_{L^2}^2 - \|\mathbf{H}_h^n\|_{L^2}^2 \right) + \frac{1}{8\tau} \left(\|\mathbf{H}_h^{n+1} - \mathbf{H}_h^n\|_{L^2}^2 - \|\mathbf{H}_h^n - \mathbf{H}_h^{n-1}\|_{L^2}^2 \right) \\
& + \frac{1}{8\tau} \|\mathbf{H}_h^{n+1} - 2\mathbf{H}_h^n + \mathbf{H}_h^{n-1}\|_{L^2}^2.
\end{aligned}$$

In turn, a substitution of $\mathbf{l}_h = \mathbf{u}_h^{n+1}$ in (3.3) yields

$$\frac{1}{2\tau} \left(\|\mathbf{u}_h^{n+1}\|_{L^2}^2 - \|\widehat{\mathbf{u}}_h^{n+1}\|_{L^2}^2 + \|\mathbf{u}_h^{n+1} - \widehat{\mathbf{u}}_h^{n+1}\|_{L^2}^2 \right) = 0, \tag{3.13}$$

where we have used the divergence-free condition (3.4) for q_h being p_h^{n+1}, p_h^n .

Next, we choose $\mathbf{l}_h = \nabla_h p_h^n$ in (3.3) and obtain

$$-\left(\nabla \cdot \widehat{\mathbf{u}}_h^{n+1}, p_h^n \right) = \frac{\tau}{4} \left(\|\nabla_h p_h^{n+1}\|_{L^2}^2 - \|\nabla_h p_h^n\|_{L^2}^2 - \|\nabla_h (p_h^{n+1} - p_h^n)\|_{L^2}^2 \right). \tag{3.14}$$

Furthermore, we get the following result from (3.9)

$$\frac{1}{4} \|\nabla_h (p_h^{n+1} - p_h^n)\|_{L^2}^2 = \frac{1}{\tau^2} \|\mathbf{u}_h^{n+1} - \widehat{\mathbf{u}}_h^{n+1}\|_{L^2}^2. \tag{3.15}$$

Summing up (3.11)–(3.15) leads to

$$\begin{aligned}
& \frac{\mu}{2\tau} \left(\|\mathbf{H}_h^{n+1}\|_{L^2}^2 - \|\mathbf{H}_h^n\|_{L^2}^2 \right) + \frac{\mu}{8\tau} \left(\|\mathbf{H}_h^{n+1} - \mathbf{H}_h^n\|_{L^2}^2 - \|\mathbf{H}_h^n - \mathbf{H}_h^{n-1}\|_{L^2}^2 \right) \\
& + \frac{\mu}{8\tau} \|\mathbf{H}_h^{n+1} - 2\mathbf{H}_h^n + \mathbf{H}_h^{n-1}\|_{L^2}^2 + \sigma^{-1} \|\nabla \times \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}}\|_{L^2}^2 + \sigma^{-1} \|\nabla \cdot \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}}\|_{L^2}^2
\end{aligned}$$

$$\begin{aligned}
 & + \frac{1}{2\tau} \left(\|\mathbf{u}_h^{n+1}\|_{L^2}^2 - \|\mathbf{u}_h^n\|_{L^2}^2 \right) + \nu \|\nabla \bar{\mathbf{u}}_h^{n+\frac{1}{2}}\|_{L^2}^2 + \frac{\tau}{8} \left(\|\nabla_h p_h^{n+1}\|_{L^2}^2 - \|\nabla_h p_h^n\|_{L^2}^2 \right) \\
 & \leq \sigma^{-1} \left(\nabla \times \mathbf{J}^{n+\frac{1}{2}}, \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \right) + \left(\mathbf{f}^{n+\frac{1}{2}}, \bar{\mathbf{u}}_h^{n+\frac{1}{2}} \right).
 \end{aligned} \tag{3.16}$$

For the right-hand side of (3.16), we can easily see that

$$\begin{aligned}
 \sigma^{-1} \left(\nabla \times \mathbf{J}^{n+\frac{1}{2}}, \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \right) & = \sigma^{-1} \left(\mathbf{J}^{n+\frac{1}{2}}, \nabla \times \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \right) \\
 & \leq \frac{1}{4\sigma} \|\mathbf{J}^{n+\frac{1}{2}}\|_{L^2}^2 + \sigma^{-1} \|\nabla \times \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}}\|_{L^2}^2,
 \end{aligned}$$

and

$$\left(\mathbf{f}^{n+\frac{1}{2}}, \bar{\mathbf{u}}_h^{n+\frac{1}{2}} \right) \leq \frac{1}{4\varepsilon} \|\mathbf{f}^{n+\frac{1}{2}}\|_{L^2}^2 + \varepsilon \|\bar{\mathbf{u}}_h^{n+\frac{1}{2}}\|_{L^2}^2 \leq \frac{1}{4\varepsilon} \|\mathbf{f}^{n+\frac{1}{2}}\|_{L^2}^2 + \varepsilon \|\nabla \bar{\mathbf{u}}_h^{n+\frac{1}{2}}\|_{L^2}^2,$$

where ε is an arbitrarily small constant. Substituting the above estimates into (3.16), we get the desired result (3.10) immediately. This completes the proof of Theorem 3.6. \square

4. ERROR ESTIMATES

We present the proof of the existence and uniqueness of numerical solution and the error estimates (3.6) and (3.7) in Section 4.

4.1. Preliminary results

We introduce several types of projections. Let $P_h : L^2(\Omega) \rightarrow M_h$ denote the L^2 projection which satisfies

$$(v - P_h v, q_h) = 0, \quad v \in L^2(\Omega), \quad \forall q_h \in M_h. \tag{4.1}$$

For the sake of brevity, if v is a vector function in $L^2(\Omega)$, we use $\mathbf{P}_h v$ to denote the L^2 projection of the vector function v onto \mathbf{X}_h . Furthermore, let $(\mathbf{R}_h \mathbf{u}, R_h p)$ denote the Stokes projection of $(\mathbf{u}, p) \in \mathbf{H}_0^1(\Omega) \times L^2(\Omega)/\mathbb{R}$ satisfying

$$\nu(\nabla(\mathbf{u} - \mathbf{R}_h \mathbf{u}), \nabla \mathbf{v}_h) - (p - R_h p, \nabla \cdot \mathbf{v}_h) = 0, \quad \forall \mathbf{v}_h \in \mathbf{X}_h, \tag{4.2}$$

$$(\nabla \cdot (\mathbf{u} - \mathbf{R}_h \mathbf{u}), q_h) = 0, \quad \forall q_h \in M_h. \tag{4.3}$$

We also introduce the Maxwell projection operator $\Pi_h : \mathring{\mathbf{H}}^1(\Omega) \rightarrow \mathbf{S}_h$, by

$$(\nabla \times (\mathbf{H} - \Pi_h \mathbf{H}), \nabla \times \mathbf{w}_h) + (\nabla \cdot (\mathbf{H} - \Pi_h \mathbf{H}), \nabla \cdot \mathbf{w}_h) = 0, \quad \mathbf{H} \in \mathring{\mathbf{H}}^1(\Omega), \forall \mathbf{w}_h \in \mathbf{S}_h. \tag{4.4}$$

For the above projections, the following estimates are recalled [11, 37].

Lemma 4.1. *The following estimates are valid for the L^2 projection, Stokes projection, and Maxwell projection:*

$$\|P_h v\|_{L^s} \leq C \|v\|_{L^s}, \quad \forall v \in L^2(\Omega), \tag{4.5}$$

$$\|P_h v\|_{H^1} \leq C \|v\|_{H^1}, \quad \forall v \in H_0^1(\Omega), \tag{4.6}$$

$$\|v - P_h v\|_{L^2} \leq Ch^{\ell+1} \|v\|_{H^{\ell+1}}, \tag{4.7}$$

for $m = 0, 1, 0 \leq \ell \leq r, 1 \leq s \leq \infty$, and

$$\|\mathbf{R}_h \mathbf{u}\|_{W^{1,s}} + \|R_h p\|_{L^s} \leq C (\|\mathbf{u}\|_{W^{1,s}} + \|p\|_{L^s}), \tag{4.8}$$

$$\|\mathbf{u} - \mathbf{R}_h \mathbf{u}\|_{L^s} + h \|\mathbf{u} - \mathbf{R}_h \mathbf{u}\|_{W^{1,s}} \leq Ch^{\ell+1} (\|\mathbf{u}\|_{W^{\ell+1,s}} + \|p\|_{W^{\ell,s}}), \tag{4.9}$$

$$\|p - R_h p\|_{L^s} \leq Ch^\ell (\|\mathbf{u}\|_{W^{\ell+1,s}} + \|p\|_{W^{\ell,s}}), \tag{4.10}$$

$$\|\partial_t(\mathbf{u} - \mathbf{R}_h\mathbf{u})\|_{L^s} + h\|\partial_t(p - R_hp)\|_{L^s} \leq Ch^{\ell+1}(\|\partial_t\mathbf{u}\|_{W^{\ell+1,s}} + \|\partial_t p\|_{W^{\ell,s}}), \quad (4.11)$$

for $0 \leq \ell \leq r$, $1 < s < \infty$, and

$$\|\mathbf{H} - \Pi_h\mathbf{H}\|_{L^2} + h\|\mathbf{H} - \Pi_h\mathbf{H}\|_{H^1} \leq Ch^{\ell+1}\|\mathbf{H}\|_{H^{\ell+1}}, \quad (4.12)$$

for $0 \leq \ell \leq r$, where C is a positive constant independent of h .

Next, we recall two lemmas that will be frequently used in this paper.

Lemma 4.2 ([4]). *Given v_h in the finite element spaces \mathbf{X}_h , M_h , or \mathbf{S}_h , the following inverse inequality holds*

$$\|v_h\|_{W^{m,s}} \leq Ch^{n-m+\frac{d}{s}-\frac{d}{q}}\|v_h\|_{W^{n,q}}, \quad (4.13)$$

for $0 \leq n \leq m \leq 1$, $1 \leq q \leq s \leq \infty$, where d denotes the dimension of the space and C is a positive constant independent of h .

Lemma 4.3. *The discrete gradient operator $\nabla_h : M_h \rightarrow \mathbf{X}_h$ (defined in (3.8)) satisfies the following estimates*

$$\|\nabla_h q_h\|_{L^2} \leq Ch^{-1}\|q_h\|_{L^2}, \quad (4.14)$$

$$\|\nabla_h q_h\|_{L^3} \leq Ch^{-1}\|q_h\|_{L^3}, \quad (4.15)$$

for any $q_h \in M_h$, where C is a positive constant independent of h .

Proof. The estimate (4.14) follows immediately by substituting $\mathbf{v}_h = \nabla_h q_h$ into (3.8) and inverse inequality (4.13).

It remains to prove (4.15). Given $q_h \in M_h$, it is easy to see that

$$\begin{aligned} (\nabla_h q_h, \mathbf{v}) &= (\nabla_h q_h, P_h\mathbf{v}) = -(q_h, \nabla \cdot P_h\mathbf{v}) \leq \|q_h\|_{L^3}\|\nabla \cdot P_h\mathbf{v}\|_{L^{\frac{3}{2}}} \\ &\leq C\|q_h\|_{L^3}h^{-1}\|P_h\mathbf{v}\|_{L^{\frac{3}{2}}} \leq Ch^{-1}\|q_h\|_{L^3}\|\mathbf{v}\|_{L^{\frac{3}{2}}}, \end{aligned}$$

for all $\mathbf{v} \in L^{\frac{3}{2}}(\Omega)$. Here, P_h is the L^2 projection, which has a bounded extension to $L^p(\Omega)$ for $1 \leq p \leq \infty$, with a bound independent of h ; see Lemma 6.1 of [37]. Then, using the duality between $L^3(\Omega)$ and $L^{\frac{3}{2}}(\Omega)$, it is straightforward to derive (4.15). The proof of Lemma 4.3 is completed. \square

4.2. Error equations

To establish error estimates for the scheme (3.1)–(3.4), we introduce an intermediate function $\widehat{\mathbf{R}_h\mathbf{u}^{n+1}} \in \mathbf{X}_h$, defined as

$$\left(\frac{\mathbf{R}_h\mathbf{u}^{n+1} - \widehat{\mathbf{R}_h\mathbf{u}^{n+1}}}{\tau}, \mathbf{l}_h \right) - \frac{1}{2}(R_hp^{n+1} - R_hp^n, \nabla \cdot \mathbf{l}_h) = 0, \quad \forall \mathbf{l}_h \in \mathbf{X}_h, \quad (4.16)$$

$$\text{or equivalently, } \frac{\mathbf{R}_h\mathbf{u}^{n+1} - \widehat{\mathbf{R}_h\mathbf{u}^{n+1}}}{\tau} = -\frac{1}{2}\nabla_h(R_hp^{n+1} - R_hp^n). \quad (4.17)$$

With the intermediate function defined above and the projections introduced in the previous subsection, the MHD system (1.1)–(1.3) can be rewritten as follows:

$$\begin{aligned} \mu \left(\frac{\Pi_h\mathbf{H}^{n+1} - \Pi_h\mathbf{H}^n}{\tau}, \mathbf{w}_h \right) + \sigma^{-1} \left(\nabla \times \Pi_h\widetilde{\mathbf{H}}^{n+\frac{1}{2}}, \nabla \times \mathbf{w}_h \right) + \sigma^{-1} \left(\nabla \cdot \Pi_h\widetilde{\mathbf{H}}^{n+\frac{1}{2}}, \nabla \cdot \mathbf{w}_h \right) \\ - \mu \left(\overline{\mathbf{u}}^{n+\frac{1}{2}} \times \widetilde{\mathbf{H}}^{n+\frac{1}{2}}, \nabla \times \mathbf{w}_h \right) = \sigma^{-1} \left(\nabla \times \mathbf{J}^{n+\frac{1}{2}}, \mathbf{w}_h \right) + R_{\mathbf{H}}^{n+1}(\mathbf{w}_h), \end{aligned} \quad (4.18)$$

$$\begin{aligned}
& \left(\frac{\widehat{\mathbf{R}_h \mathbf{u}^{n+1}} - \mathbf{R}_h \mathbf{u}^n}{\tau}, \mathbf{v}_h \right) + \nu \left(\nabla \left(\frac{\widehat{\mathbf{R}_h \mathbf{u}^{n+1}} + \mathbf{R}_h \mathbf{u}^n}{2} \right), \nabla \mathbf{v}_h \right) + b \left(\tilde{\mathbf{u}}^{n+\frac{1}{2}}, \bar{\mathbf{u}}^{n+\frac{1}{2}}, \mathbf{v}_h \right) \\
& \quad - (R_h p^n, \nabla \cdot \mathbf{v}_h) + \mu \left(\widetilde{\mathbf{H}}^{n+\frac{1}{2}} \times \left(\nabla \times \widetilde{\mathbf{H}}^{n+\frac{1}{2}} \right), \mathbf{v}_h \right) \\
& = \left(\mathbf{f}^{n+\frac{1}{2}}, \mathbf{v}_h \right) + \left(\frac{\widehat{\mathbf{R}_h \mathbf{u}^{n+1}} - \mathbf{R}_h \mathbf{u}^{n+1}}{\tau}, \mathbf{v}_h \right) + \nu \left(\nabla \left(\frac{\widehat{\mathbf{R}_h \mathbf{u}^{n+1}} - \mathbf{R}_h \mathbf{u}^{n+1}}{2} \right), \nabla \mathbf{v}_h \right) \\
& \quad - \left(R_h p^n - \frac{R_h p^{n+1} + R_h p^n}{2}, \nabla \cdot \mathbf{v}_h \right) + R_{\mathbf{u}}^{n+1}(\mathbf{v}_h), \tag{4.19}
\end{aligned}$$

$$(\nabla \cdot \mathbf{u}^{n+1}, q_h) = 0, \tag{4.20}$$

for any $(\mathbf{w}_h, \mathbf{v}_h, q_h) \in (\mathbf{S}_h, \mathbf{X}_h, M_h)$ and $n = 1, 2, \dots, N-1$, where we denote $\hat{\mathbf{u}}^{n+1} := \mathbf{u}^{n+1}$, $R_{\mathbf{H}}^{n+1}(\mathbf{w}_h)$ and $R_{\mathbf{u}}^{n+1}(\mathbf{v}_h)$ stand for the truncation errors satisfying

$$\begin{aligned}
R_{\mathbf{H}}^{n+1}(\mathbf{w}_h) &= \mu \left(\frac{\Pi_h \mathbf{H}^{n+1} - \Pi_h \mathbf{H}^n}{\tau} - \partial_t \mathbf{H}^{n+\frac{1}{2}}, \mathbf{w}_h \right) + \sigma^{-1} \left(\nabla \times \left(\Pi_h \widetilde{\mathbf{H}}^{n+\frac{1}{2}} - \mathbf{H}^{n+\frac{1}{2}} \right), \nabla \times \mathbf{w}_h \right) \\
& \quad + \sigma^{-1} \left(\nabla \cdot \left(\Pi_h \widetilde{\mathbf{H}}^{n+\frac{1}{2}} - \mathbf{H}^{n+\frac{1}{2}} \right), \nabla \cdot \mathbf{w}_h \right) - \mu \left(\bar{\mathbf{u}}^{n+\frac{1}{2}} \times \widetilde{\mathbf{H}}^{n+\frac{1}{2}} - \mathbf{u}^{n+\frac{1}{2}} \times \mathbf{H}^{n+\frac{1}{2}}, \nabla \times \mathbf{w}_h \right), \tag{4.21}
\end{aligned}$$

$$\begin{aligned}
R_{\mathbf{u}}^{n+1}(\mathbf{v}_h) &= \left(\frac{\mathbf{R}_h \mathbf{u}^{n+1} - \mathbf{R}_h \mathbf{u}^n}{\tau} - \partial_t \mathbf{u}^{n+\frac{1}{2}}, \mathbf{v}_h \right) + \nu \left(\nabla \left(\frac{\mathbf{R}_h \mathbf{u}^{n+1} + \mathbf{R}_h \mathbf{u}^n}{2} - \mathbf{u}^{n+\frac{1}{2}} \right), \nabla \mathbf{v}_h \right) \\
& \quad + \left(b \left(\tilde{\mathbf{u}}^{n+\frac{1}{2}}, \bar{\mathbf{u}}^{n+\frac{1}{2}}, \mathbf{v}_h \right) - b \left(\mathbf{u}^{n+\frac{1}{2}}, \mathbf{u}^{n+\frac{1}{2}}, \mathbf{v}_h \right) \right) - \left(\frac{R_h p^{n+1} + R_h p^n}{2} - p^{n+\frac{1}{2}}, \nabla \cdot \mathbf{v}_h \right) \\
& \quad + \mu \left(\widetilde{\mathbf{H}}^{n+\frac{1}{2}} \times \left(\nabla \times \widetilde{\mathbf{H}}^{n+\frac{1}{2}} \right) - \mathbf{H}^{n+\frac{1}{2}} \times \left(\nabla \times \mathbf{H}^{n+\frac{1}{2}} \right), \mathbf{v}_h \right). \tag{4.22}
\end{aligned}$$

Utilizing the projection error estimates presented in the previous subsection, we only need to estimate the following error functions

$$\begin{aligned}
e_{\mathbf{H}}^n &= \Pi_h \mathbf{H}^n - \mathbf{H}_h^n, & e_{\mathbf{u}}^n &= \mathbf{R}_h \mathbf{u}^n - \mathbf{u}_h^n, \\
\hat{e}_{\mathbf{u}}^n &= \widehat{\mathbf{R}_h \mathbf{u}^n} - \hat{\mathbf{u}}_h^n, & e_p^n &= R_h p^n - p_h^n,
\end{aligned}$$

for $n = 1, 2, \dots, N$. From the system (4.16)–(4.20) and the fully discrete numerical scheme (3.1)–(3.4), we observe that the error functions $(e_{\mathbf{H}}^n, e_{\mathbf{u}}^n, \hat{e}_{\mathbf{u}}^n, e_p^n)$ satisfy the following equations:

$$\begin{aligned}
& \mu \left(\frac{e_{\mathbf{H}}^{n+1} - e_{\mathbf{H}}^n}{\tau}, \mathbf{w}_h \right) + \sigma^{-1} \left(\nabla \times \tilde{e}_{\mathbf{H}}^{n+\frac{1}{2}}, \nabla \times \mathbf{w}_h \right) + \sigma^{-1} \left(\nabla \cdot \tilde{e}_{\mathbf{H}}^{n+\frac{1}{2}}, \nabla \cdot \mathbf{w}_h \right) \\
& \quad = \mu \left\{ \left(\bar{\mathbf{u}}^{n+\frac{1}{2}} \times \widetilde{\mathbf{H}}^{n+\frac{1}{2}}, \nabla \times \mathbf{w}_h \right) - \left(\bar{\mathbf{u}}_h^{n+\frac{1}{2}} \times \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}}, \nabla \times \mathbf{w}_h \right) \right\} + R_{\mathbf{H}}^{n+1}(\mathbf{w}_h), \tag{4.23}
\end{aligned}$$

$$\begin{aligned}
& \left(\frac{\hat{e}_{\mathbf{u}}^{n+1} - e_{\mathbf{u}}^n}{\tau}, \mathbf{v}_h \right) + \nu \left(\nabla \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}}, \nabla \mathbf{v}_h \right) - (e_p^n, \nabla \cdot \mathbf{v}_h) \\
& \quad = \left(\frac{\widehat{\mathbf{R}_h \mathbf{u}^{n+1}} - \mathbf{R}_h \mathbf{u}^{n+1}}{\tau}, \mathbf{v}_h \right) + \nu \left(\nabla \left(\frac{\widehat{\mathbf{R}_h \mathbf{u}^{n+1}} - \mathbf{R}_h \mathbf{u}^{n+1}}{2} \right), \nabla \mathbf{v}_h \right) \\
& \quad \quad - \left(R_h p^n - \frac{R_h p^{n+1} + R_h p^n}{2}, \nabla \cdot \mathbf{v}_h \right) - \left\{ b \left(\tilde{\mathbf{u}}^{n+\frac{1}{2}}, \bar{\mathbf{u}}^{n+\frac{1}{2}}, \mathbf{v}_h \right) - b \left(\tilde{\mathbf{u}}_h^{n+\frac{1}{2}}, \bar{\mathbf{u}}_h^{n+\frac{1}{2}}, \mathbf{v}_h \right) \right\} \\
& \quad \quad - \mu \left\{ \left(\widetilde{\mathbf{H}}^{n+\frac{1}{2}} \times \left(\nabla \times \widetilde{\mathbf{H}}^{n+\frac{1}{2}} \right), \mathbf{v}_h \right) - \left(\widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \times \left(\nabla \times \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \right), \mathbf{v}_h \right) \right\} + R_{\mathbf{u}}^{n+1}(\mathbf{v}_h), \tag{4.24}
\end{aligned}$$

$$\left(\frac{e_{\mathbf{u}}^{n+1} - \widehat{e}_{\mathbf{u}}^{n+1}}{\tau}, \mathbf{l}_h \right) - \frac{1}{2} (e_p^{n+1} - e_p^n, \nabla \cdot \mathbf{l}_h) = 0, \quad (4.25)$$

$$(\nabla \cdot e_{\mathbf{u}}^{n+1}, q_h) = 0, \quad (4.26)$$

for any $(\mathbf{w}_h, \mathbf{v}_h, \mathbf{l}_h, q_h) \in (\mathcal{S}_h, \mathbf{X}_h, \mathbf{X}_h, M_h)$ and $n = 1, 2, \dots, N-1$.

4.3. Proof of Theorem 3.1

In this subsection, we present a detailed proof of Theorem 3.1. The following lemma will be used in the analysis.

Lemma 4.4. *Under the regularity assumption (3.5), the Stokes projection defined in (4.2) and (4.3) satisfies the following estimates:*

$$\|\nabla_h R_h \partial_t p\|_{L^3} \leq C, \quad (4.27)$$

$$\|\nabla(\nabla_h R_h \partial_t p)\|_{L^2} \leq C \ell_h, \quad (4.28)$$

where C is a positive constant independent of h ; $\ell_h = 1$ if $\nabla p|_{\partial\Omega} = \mathbf{0}$, otherwise $\ell_h = h^{-\frac{1}{2}}$.

Proof. By the regularity assumption (3.5) and the L^3 stability estimate of the L^2 projection, i.e., (4.5), we see that

$$\|\mathbf{P}_h \nabla \partial_t p\|_{L^3} \leq C \|\nabla \partial_t p\|_{L^3} \leq C. \quad (4.29)$$

Since

$$\begin{aligned} (\mathbf{v}_h, \mathbf{P}_h \nabla \partial_t p - \nabla_h P_h \partial_t p) &= (\mathbf{v}_h, \nabla \partial_t p) + (\nabla \cdot \mathbf{v}_h, P_h \partial_t p) \\ &= -(\nabla \cdot \mathbf{v}_h, \partial_t p) + (\nabla \cdot \mathbf{v}_h, P_h \partial_t p) \\ &\leq \|\nabla \cdot \mathbf{v}_h\|_{L^{\frac{3}{2}}} \|\partial_t p - P_h \partial_t p\|_{L^3} \\ &\leq Ch^{-1} \|\mathbf{v}_h\|_{L^{\frac{3}{2}}} h \|\partial_t p\|_{W^{1,3}} \\ &\leq C \|\mathbf{v}_h\|_{L^{\frac{3}{2}}} \|\partial_t p\|_{W^{1,3}}, \end{aligned}$$

for any $\mathbf{v}_h \in \mathbf{X}_h$, by the duality between $L^{\frac{3}{2}}$ and L^3 , we conclude that

$$\|\mathbf{P}_h \nabla \partial_t p - \nabla_h P_h \partial_t p\|_{L^3} \leq C. \quad (4.30)$$

Consequently, with the help of the inverse inequality (4.15), we obtain

$$\begin{aligned} \|\nabla_h R_h \partial_t p\|_{L^3} &\leq \|\nabla_h R_h \partial_t p - \nabla_h P_h \partial_t p\|_{L^3} + \|\nabla_h P_h \partial_t p - \mathbf{P}_h \nabla \partial_t p\|_{L^3} + \|\mathbf{P}_h \nabla \partial_t p\|_{L^3} \\ &\leq \|\nabla_h R_h \partial_t p - \nabla_h P_h \partial_t p\|_{L^3} + C \\ &\leq Ch^{-1} \|R_h \partial_t p - P_h \partial_t p\|_{L^3} + C \\ &\leq Ch^{-1} \|R_h \partial_t p - \partial_t p\|_{L^3} + Ch^{-1} \|\partial_t p - P_h \partial_t p\|_{L^3} + C \\ &\leq Ch^{-1} h^2 + Ch^{-1} h^2 + C \leq C, \end{aligned}$$

in which (4.7) and the projection estimate (4.11) have been used in the second last inequality.

Inequality (4.28) could be proved in a similar manner. If $\nabla p|_{\partial\Omega} = \mathbf{0}$, by the regularity assumption (3.5) and the H^1 stability estimate of the L^2 projection, we have

$$\|\nabla \mathbf{P}_h \nabla \partial_t p\|_{L^2} \leq \|\mathbf{P}_h \nabla \partial_t p\|_{H^1} \leq C \|\nabla \partial_t p\|_{H^1} \leq C. \quad (4.31)$$

If $\nabla p|_{\partial\Omega} \neq \mathbf{0}$, we estimate $\|\nabla \mathbf{P}_h \nabla \partial_t p\|_{L^2}$ in another way. Let χ be a smooth cut-off function such that

$$\begin{cases} \chi = 1 & \text{on } \partial\Omega, \\ \chi = 0 & \text{at points } x \text{ such that } \text{dist}(x, \partial\Omega) \geq h, \\ 0 \leq \chi \leq 1 & \text{in } \Omega, \\ \|\nabla^k \chi\|_{L^\infty} \leq Ch^{-k} & \text{in } \Omega. \end{cases}$$

Then we let $\mathbf{g} = \nabla \partial_t p$ and get

$$\|\nabla \mathbf{P}_h \mathbf{g}\|_{L^2} \leq \|\nabla \mathbf{P}_h (\mathbf{g} - \chi \mathbf{g})\|_{L^2} + \|\nabla \mathbf{P}_h \chi \mathbf{g}\|_{L^2}.$$

Since $\mathbf{g} - \chi \mathbf{g} = \mathbf{0}$ on $\partial\Omega$, it follows that

$$\begin{aligned} \|\nabla \mathbf{P}_h (\mathbf{g} - \chi \mathbf{g})\|_{L^2} &\leq C \|\nabla (\mathbf{g} - \chi \mathbf{g})\|_{L^2} \leq C \|\nabla \mathbf{g}\|_{L^2} + C \|\nabla \chi \otimes \mathbf{g}\|_{L^2} \\ &\leq C \|\nabla \mathbf{g}\|_{L^2} + Ch^{-1} \left(\int_{\text{dist}(x, \partial\Omega) \leq h} |\mathbf{g}|^2 dx \right)^{\frac{1}{2}} \leq Ch^{-\frac{1}{2}}. \end{aligned}$$

By the inverse inequality, we have

$$\|\nabla \mathbf{P}_h \chi \mathbf{g}\|_{L^2} \leq Ch^{-1} \|\mathbf{P}_h \chi \mathbf{g}\|_{L^2} \leq Ch^{-1} \|\chi \mathbf{g}\|_{L^2} \leq Ch^{-1} \left(\int_{\text{dist}(x, \partial\Omega) \leq h} |\mathbf{g}|^2 dx \right)^{\frac{1}{2}} \leq Ch^{-\frac{1}{2}}.$$

Therefore,

$$\|\nabla \mathbf{P}_h \mathbf{g}\|_{L^2} \leq Ch^{-\frac{1}{2}}, \tag{4.32}$$

which together with (4.31) implies

$$\|\nabla \mathbf{P}_h \nabla \partial_t p\|_{L^2} \leq Cl_h. \tag{4.33}$$

Using similar techniques in the derivation of (4.30), we get

$$\|\mathbf{P}_h \nabla \partial_t p - \nabla_h \mathbf{P}_h \partial_t p\|_{L^2} \leq Ch. \tag{4.34}$$

By the inverse inequalities (4.13) and (4.14), it can be shown that

$$\begin{aligned} \|\nabla (\nabla_h R_h \partial_t p - \mathbf{P}_h \nabla \partial_t p)\|_{L^2} &\leq Ch^{-1} \|\nabla_h R_h \partial_t p - \mathbf{P}_h \nabla \partial_t p\|_{L^2} \\ &\leq Ch^{-1} \|\nabla_h R_h \partial_t p - \nabla_h \mathbf{P}_h \partial_t p\|_{L^2} + Ch^{-1} \|\nabla_h \mathbf{P}_h \partial_t p - \mathbf{P}_h \nabla \partial_t p\|_{L^2} \\ &\leq Ch^{-2} \|R_h \partial_t p - \mathbf{P}_h \partial_t p\|_{L^2} + Ch^{-1} h \\ &\leq Ch^{-2} \|R_h \partial_t p - \partial_t p\|_{L^2} + Ch^{-2} \|\partial_t p - \mathbf{P}_h \partial_t p\|_{L^2} + C \\ &\leq Ch^{-2} h^2 + Ch^{-2} h^2 + C \leq C, \end{aligned} \tag{4.35}$$

in which (4.7) and (4.11) have been used again in the second to last inequality. Finally, by the triangle inequality and (4.31)–(4.35), the estimate (4.28) follows immediately. \square

Now we proceed with the proof of Theorem 3.1.

Proof of Theorem 3.1. By Theorem 3.6, the existence and uniqueness of numerical solution $(\mathbf{H}_h^n, \mathbf{u}_h^n, p_h^n)$, $n = 2, 3, \dots, N$, follows immediately since the scheme (3.1)–(3.4) is linearized and the corresponding homogeneous equations only admit zero solutions.

In the following, we present the analysis of the error equations (4.23)–(4.26) and then establish the error estimates given in Theorem 3.1. First of all, we make the following induction assumption for the error functions at the previous time steps:

$$\|e_{\mathbf{H}}^m\|_{L^2} + \|e_{\mathbf{u}}^m\|_{L^2} \leq C_0^*(\ell_h \tau^2 + h^2), \quad (4.36)$$

for $m \leq n$. Such an induction assumption will be recovered by the error estimate at the next time step t_{n+1} .

For $m = 0, 1$, (4.36) follows from Remark 3.5 immediately. The induction assumption (4.36) (for $m \leq n$) yields

$$\begin{aligned} \|\mathbf{H}_h^m\|_{W^{1,3}} &\leq \|I_h \mathbf{H}^m\|_{W^{1,3}} + \|I_h \mathbf{H}^m - \Pi_h \mathbf{H}^m\|_{W^{1,3}} + \|e_{\mathbf{H}}^m\|_{W^{1,3}} \\ &\leq C\|\mathbf{H}^m\|_{W^{1,3}} + Ch^{-\frac{d}{6}}\|I_h \mathbf{H}^m - \Pi_h \mathbf{H}^m\|_{H^1} + Ch^{-1-\frac{d}{6}}\|e_{\mathbf{H}}^m\|_{L^2} \\ &\leq C\|\mathbf{H}^m\|_{W^{1,3}} + Ch^{-\frac{d}{6}}\|I_h \mathbf{H}^m - \mathbf{H}^m\|_{H^1} + Ch^{-\frac{d}{6}}\|\mathbf{H}^m - \Pi_h \mathbf{H}^m\|_{H^1} \\ &\quad + Ch^{-1-\frac{d}{6}}C_0^*(\ell_h \tau^2 + h^2) \\ &\leq C\|\mathbf{H}^m\|_{W^{1,3}} + Ch^{-\frac{d}{6}}h^2 + Ch^{-\frac{d}{6}}h^2 + Ch^{-1-\frac{d}{6}}C_0^*(\ell_h \tau^2 + h^2) \\ &\leq K + 1, \end{aligned} \quad (4.37)$$

$$\begin{aligned} \|\mathbf{u}_h^m\|_{L^\infty} &\leq \|\mathbf{R}_h \mathbf{u}^m\|_{L^\infty} + \|e_{\mathbf{u}}^m\|_{L^\infty} \\ &\leq \|\mathbf{u}^m\|_{W^{1,3}} + Ch^{-\frac{d}{2}}\|e_{\mathbf{u}}^m\|_{L^2} \quad (\text{by (4.8)}) \\ &\leq \|\mathbf{u}^m\|_{W^{1,3}} + Ch^{-\frac{d}{2}}C_0^*(\ell_h \tau^2 + h^2) \\ &\leq K + 1, \end{aligned} \quad (4.38)$$

for $\tau \leq \frac{h}{\sqrt{2CC_0^*}}$ and $h < h_0$, where $d = 2, 3$, denotes the dimension of Ω and h_0 is a small positive constant. Here, I_h denotes the standard Lagrange interpolation and its $W^{1,3}$ stability estimate has been used. Subsequently, we will establish the error estimate at $m = n + 1$ and recover (4.36).

Step 1: Estimate of (4.23). Taking $\mathbf{w}_h = \check{e}_{\mathbf{H}}^{n+\frac{1}{2}}$ into (4.23) yields

$$\begin{aligned} &\frac{\mu}{2\tau} \left(\|e_{\mathbf{H}}^{n+1}\|_{L^2}^2 - \|e_{\mathbf{H}}^n\|_{L^2}^2 \right) + \frac{\mu}{8\tau} \left(\|e_{\mathbf{H}}^{n+1} - e_{\mathbf{H}}^n\|_{L^2}^2 - \|e_{\mathbf{H}}^n - e_{\mathbf{H}}^{n-1}\|_{L^2}^2 \right) \\ &\quad + \sigma^{-1} \|\nabla \times \check{e}_{\mathbf{H}}^{n+\frac{1}{2}}\|_{L^2}^2 + \sigma^{-1} \|\nabla \cdot \check{e}_{\mathbf{H}}^{n+\frac{1}{2}}\|_{L^2}^2 \\ &\leq \mu \left\{ \left(\bar{\mathbf{u}}^{n+\frac{1}{2}} \times \widetilde{\mathbf{H}}^{n+\frac{1}{2}}, \nabla \times \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right) - \left(\bar{\mathbf{u}}_h^{n+\frac{1}{2}} \times \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}}, \nabla \times \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right) \right\} + R_{\mathbf{H}}^{n+1} \left(\check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right), \end{aligned} \quad (4.39)$$

where we have used the identity

$$\begin{aligned} \left(\frac{e_{\mathbf{H}}^{n+1} - e_{\mathbf{H}}^n}{\tau}, \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right) &= \frac{1}{2\tau} \left(\|e_{\mathbf{H}}^{n+1}\|_{L^2}^2 - \|e_{\mathbf{H}}^n\|_{L^2}^2 \right) + \frac{1}{8\tau} \left(\|e_{\mathbf{H}}^{n+1} - e_{\mathbf{H}}^n\|_{L^2}^2 - \|e_{\mathbf{H}}^n - e_{\mathbf{H}}^{n-1}\|_{L^2}^2 \right) \\ &\quad + \frac{1}{8\tau} \|e_{\mathbf{H}}^{n+1} - 2e_{\mathbf{H}}^n + e_{\mathbf{H}}^{n-1}\|_{L^2}^2. \end{aligned} \quad (4.40)$$

By (3.5) and (4.12), it can be shown that

$$R_{\mathbf{H}}^{n+1} \left(\check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right) \leq C(\tau^2 + h^{r+1})^2 + C\|\check{e}_{\mathbf{H}}^{n+\frac{1}{2}}\|_{L^2}^2 + \frac{1}{2\sigma} \|\nabla \times \check{e}_{\mathbf{H}}^{n+\frac{1}{2}}\|_{L^2}^2 + \frac{1}{2\sigma} \|\nabla \cdot \check{e}_{\mathbf{H}}^{n+\frac{1}{2}}\|_{L^2}^2.$$

Noticing that $\hat{\mathbf{u}}^{n+1} := \mathbf{u}^{n+1}$ and (4.17), we obtain

$$\mu \left\{ \left(\bar{\mathbf{u}}^{n+\frac{1}{2}} \times \widetilde{\mathbf{H}}^{n+\frac{1}{2}}, \nabla \times \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right) - \left(\bar{\mathbf{u}}_h^{n+\frac{1}{2}} \times \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}}, \nabla \times \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right) \right\}$$

$$\begin{aligned}
 &= \mu \left(\bar{\mathbf{u}}^{n+\frac{1}{2}} \times \left(\widetilde{\mathbf{H}}^{n+\frac{1}{2}} - \Pi_h \widetilde{\mathbf{H}}^{n+\frac{1}{2}} \right), \nabla \times \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right) + \mu \left(\bar{\mathbf{u}}^{n+\frac{1}{2}} \times \check{e}_{\mathbf{H}}^{n+\frac{1}{2}}, \nabla \times \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right) \\
 &\quad + \mu \left(\left(\bar{\mathbf{u}}^{n+\frac{1}{2}} - \frac{\mathbf{R}_h \mathbf{u}^{n+1} + \mathbf{R}_h \mathbf{u}^n}{2} \right) \times \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}}, \nabla \times \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right) \\
 &\quad + \mu \left(\frac{\mathbf{R}_h \mathbf{u}^{n+1} - \widehat{\mathbf{R}_h \mathbf{u}^{n+1}}}{2} \times \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}}, \nabla \times \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right) \\
 &\quad + \mu \left(\bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \times \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}}, \nabla \times \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right) \\
 &\leq \mu \left\| \bar{\mathbf{u}}^{n+\frac{1}{2}} \right\|_{L^\infty} \left\| \widetilde{\mathbf{H}}^{n+\frac{1}{2}} - \Pi_h \widetilde{\mathbf{H}}^{n+\frac{1}{2}} \right\|_{L^2} \left\| \nabla \times \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right\|_{L^2} + \mu \left\| \bar{\mathbf{u}}^{n+\frac{1}{2}} \right\|_{L^\infty} \left\| \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right\|_{L^2} \left\| \nabla \times \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right\|_{L^2} \\
 &\quad + \mu \left\| \frac{\mathbf{u}^{n+1} + \mathbf{u}^n}{2} - \frac{\mathbf{R}_h \mathbf{u}^{n+1} + \mathbf{R}_h \mathbf{u}^n}{2} \right\|_{L^3} \left\| \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \right\|_{L^6} \left\| \nabla \times \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right\|_{L^2} \\
 &\quad + \frac{\mu\tau}{4} \left\| \nabla_h (\mathbf{R}_h p^{n+1} - \mathbf{R}_h p^n) \right\|_{L^3} \left\| \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \right\|_{L^6} \left\| \nabla \times \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right\|_{L^2} \\
 &\quad + \mu \left(\bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \times \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}}, \nabla \times \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right) \\
 &\leq Ch^{2(r+1)} + \frac{1}{4\sigma} \left\| \nabla \times \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right\|_{L^2}^2 + C \left\| \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right\|_{L^2}^2 + C\tau^4 + \mu \left(\bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \times \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}}, \nabla \times \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right),
 \end{aligned}$$

where in the last inequality we have used the projection estimates (4.9) and (4.12), (4.37), Lemma 4.4 and the following inequality:

$$\frac{\mu\tau}{4} \left\| \nabla_h (\mathbf{R}_h p^{n+1} - \mathbf{R}_h p^n) \right\|_{L^3} \leq C\tau^2.$$

With the above results, (4.39) is reduced to

$$\begin{aligned}
 &\frac{\mu}{2\tau} \left(\|e_{\mathbf{H}}^{n+1}\|_{L^2}^2 - \|e_{\mathbf{H}}^n\|_{L^2}^2 \right) + \frac{\mu}{8\tau} \left(\|e_{\mathbf{H}}^{n+1} - e_{\mathbf{H}}^n\|_{L^2}^2 - \|e_{\mathbf{H}}^n - e_{\mathbf{H}}^{n-1}\|_{L^2}^2 \right) \\
 &\quad + \frac{1}{4\sigma} \left\| \nabla \times \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right\|_{L^2}^2 + \frac{1}{4\sigma} \left\| \nabla \cdot \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right\|_{L^2}^2 \\
 &\leq C(\tau^2 + h^{r+1})^2 + C \left\| \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right\|_{L^2}^2 + C \left\| \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right\|_{L^2}^2 + \mu \left(\bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \times \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}}, \nabla \times \check{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right). \tag{4.41}
 \end{aligned}$$

Step 2: Estimate of (4.24). Taking $\mathbf{v}_h = \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} = \frac{1}{2}(\hat{e}_{\mathbf{u}}^{n+1} + e_{\mathbf{u}}^n)$ into (4.24) leads to

$$\begin{aligned}
 &\frac{1}{2\tau} \left(\|\hat{e}_{\mathbf{u}}^{n+1}\|_{L^2}^2 - \|e_{\mathbf{u}}^n\|_{L^2}^2 \right) + \nu \left\| \nabla \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right\|_{L^2}^2 - \left(e_p^n, \nabla \cdot \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right) \\
 &\leq \left(\frac{\widehat{\mathbf{R}_h \mathbf{u}^{n+1}} - \mathbf{R}_h \mathbf{u}^{n+1}}{\tau}, \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right) + \nu \left(\nabla \left(\frac{\widehat{\mathbf{R}_h \mathbf{u}^{n+1}} - \mathbf{R}_h \mathbf{u}^{n+1}}{2} \right), \nabla \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right) \\
 &\quad - \left(\mathbf{R}_h p^n - \frac{\mathbf{R}_h p^{n+1} + \mathbf{R}_h p^n}{2}, \nabla \cdot \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right) \\
 &\quad - \left\{ b \left(\widetilde{\mathbf{u}}^{n+\frac{1}{2}}, \bar{\mathbf{u}}^{n+\frac{1}{2}}, \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right) - b \left(\widetilde{\mathbf{u}}_h^{n+\frac{1}{2}}, \bar{\mathbf{u}}_h^{n+\frac{1}{2}}, \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right) \right\} \\
 &\quad - \mu \left\{ \left(\widetilde{\mathbf{H}}^{n+\frac{1}{2}} \times \left(\nabla \times \widetilde{\mathbf{H}}^{n+\frac{1}{2}} \right), \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right) - \left(\widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \times \left(\nabla \times \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \right), \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right) \right\} + R_{\mathbf{u}}^{n+1} \left(\bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right) \\
 &=: \sum_{j=1}^6 \mathcal{I}_j. \tag{4.42}
 \end{aligned}$$

In the following, we estimate \mathcal{I}_j , $j = 1, 2, \dots, 6$, respectively. By using (4.16), we have

$$\mathcal{I}_1 + \mathcal{I}_3 = - \left(\frac{R_h p^{n+1} + R_h p^n}{2} - \frac{R_h p^{n+1} + R_h p^n}{2}, \nabla \cdot \bar{e}_u^{n+\frac{1}{2}} \right) = 0.$$

By (4.17), \mathcal{I}_2 becomes

$$\begin{aligned} \mathcal{I}_2 &= \frac{\nu\tau}{4} \left(\nabla(\nabla_h(R_h p^{n+1} - R_h p^n)), \nabla \bar{e}_u^{n+\frac{1}{2}} \right) \\ &\leq C\tau^2 \|\nabla(\nabla_h(R_h p^{n+1} - R_h p^n))\|_{L^2}^2 + \varepsilon \|\nabla \bar{e}_u^{n+\frac{1}{2}}\|_{L^2}^2 \\ &\leq C\ell_h^2 \tau^4 + \varepsilon \|\nabla \bar{e}_u^{n+\frac{1}{2}}\|_{L^2}^2, \end{aligned}$$

where we have used the second result in Lemma 4.4. By the definition of $b(\mathbf{u}, \mathbf{v}, \mathbf{w})$ in (2.4), we can rewrite \mathcal{I}_4 as

$$\begin{aligned} \mathcal{I}_4 &= \frac{1}{2} \left\{ \left(\tilde{\mathbf{u}}_h^{n+\frac{1}{2}} \cdot \nabla \bar{\mathbf{u}}_h^{n+\frac{1}{2}}, \bar{e}_u^{n+\frac{1}{2}} \right) - \left(\tilde{\mathbf{u}}^{n+\frac{1}{2}} \cdot \nabla \bar{\mathbf{u}}^{n+\frac{1}{2}}, \bar{e}_u^{n+\frac{1}{2}} \right) \right\} \\ &\quad - \frac{1}{2} \left\{ \left(\tilde{\mathbf{u}}_h^{n+\frac{1}{2}} \cdot \nabla \bar{e}_u^{n+\frac{1}{2}}, \bar{\mathbf{u}}_h^{n+\frac{1}{2}} \right) - \left(\tilde{\mathbf{u}}^{n+\frac{1}{2}} \cdot \nabla \bar{e}_u^{n+\frac{1}{2}}, \bar{\mathbf{u}}^{n+\frac{1}{2}} \right) \right\} \\ &= -\frac{1}{2} \left\{ \left(\tilde{\mathbf{u}}_h^{n+\frac{1}{2}} \cdot \nabla \bar{e}_u^{n+\frac{1}{2}}, \bar{e}_u^{n+\frac{1}{2}} \right) + \left(\tilde{\mathbf{u}}_h^{n+\frac{1}{2}} \cdot \nabla \left(\bar{\mathbf{u}}^{n+\frac{1}{2}} - \overline{\mathbf{R}_h \mathbf{u}^{n+\frac{1}{2}}} \right), \bar{e}_u^{n+\frac{1}{2}} \right) \right. \\ &\quad \left. + \left(\bar{e}_u^{n+\frac{1}{2}} \cdot \nabla \bar{\mathbf{u}}^{n+\frac{1}{2}}, \bar{e}_u^{n+\frac{1}{2}} \right) + \left(\left(\tilde{\mathbf{u}}^{n+\frac{1}{2}} - \widetilde{\mathbf{R}_h \mathbf{u}^{n+\frac{1}{2}}} \right) \cdot \nabla \bar{\mathbf{u}}^{n+\frac{1}{2}}, \bar{e}_u^{n+\frac{1}{2}} \right) \right\} \\ &\quad + \frac{1}{2} \left\{ \left(\tilde{\mathbf{u}}_h^{n+\frac{1}{2}} \cdot \nabla \bar{e}_u^{n+\frac{1}{2}}, \bar{e}_u^{n+\frac{1}{2}} \right) + \left(\tilde{\mathbf{u}}_h^{n+\frac{1}{2}} \cdot \nabla \bar{e}_u^{n+\frac{1}{2}}, \bar{\mathbf{u}}^{n+\frac{1}{2}} - \overline{\mathbf{R}_h \mathbf{u}^{n+\frac{1}{2}}} \right) \right. \\ &\quad \left. + \left(\bar{e}_u^{n+\frac{1}{2}} \cdot \nabla \bar{e}_u^{n+\frac{1}{2}}, \bar{\mathbf{u}}^{n+\frac{1}{2}} \right) + \left(\left(\tilde{\mathbf{u}}^{n+\frac{1}{2}} - \widetilde{\mathbf{R}_h \mathbf{u}^{n+\frac{1}{2}}} \right) \cdot \nabla \bar{e}_u^{n+\frac{1}{2}}, \bar{\mathbf{u}}^{n+\frac{1}{2}} \right) \right\} \\ &=: \frac{1}{2} \sum_{k=1}^8 \mathcal{I}_{4,k}. \end{aligned}$$

In the estimate of \mathcal{I}_4 , the most difficult processing is the control of $\mathcal{I}_{4,2}$, for which an application of integration by parts implies that

$$\begin{aligned} \mathcal{I}_{4,2} &= \left((\nabla \cdot \tilde{\mathbf{u}}_h^{n+\frac{1}{2}}) \left(\bar{\mathbf{u}}^{n+\frac{1}{2}} - \overline{\mathbf{R}_h \mathbf{u}^{n+\frac{1}{2}}} \right), \bar{e}_u^{n+\frac{1}{2}} \right) + \left(\tilde{\mathbf{u}}_h^{n+\frac{1}{2}} \cdot \nabla \bar{e}_u^{n+\frac{1}{2}}, \bar{\mathbf{u}}^{n+\frac{1}{2}} - \overline{\mathbf{R}_h \mathbf{u}^{n+\frac{1}{2}}} \right) \\ &= \left((\nabla \cdot \widetilde{\mathbf{R}_h \mathbf{u}^{n+\frac{1}{2}}}) \left(\bar{\mathbf{u}}^{n+\frac{1}{2}} - \overline{\mathbf{R}_h \mathbf{u}^{n+\frac{1}{2}}} \right), \bar{e}_u^{n+\frac{1}{2}} \right) - \left((\nabla \cdot \tilde{e}_u^{n+\frac{1}{2}}) \left(\bar{\mathbf{u}}^{n+\frac{1}{2}} - \overline{\mathbf{R}_h \mathbf{u}^{n+\frac{1}{2}}} \right), \bar{e}_u^{n+\frac{1}{2}} \right) \\ &\quad + \left(\tilde{\mathbf{u}}_h^{n+\frac{1}{2}} \cdot \nabla \bar{e}_u^{n+\frac{1}{2}}, \bar{\mathbf{u}}^{n+\frac{1}{2}} - \overline{\mathbf{R}_h \mathbf{u}^{n+\frac{1}{2}}} \right) \\ &\leq \left\| \nabla \cdot \widetilde{\mathbf{R}_h \mathbf{u}^{n+\frac{1}{2}}} \right\|_{L^3} \left\| \bar{\mathbf{u}}^{n+\frac{1}{2}} - \overline{\mathbf{R}_h \mathbf{u}^{n+\frac{1}{2}}} \right\|_{L^2} \left\| \bar{e}_u^{n+\frac{1}{2}} \right\|_{L^6} \\ &\quad + \left\| \nabla \cdot \tilde{e}_u^{n+\frac{1}{2}} \right\|_{L^3} \left\| \bar{\mathbf{u}}^{n+\frac{1}{2}} - \overline{\mathbf{R}_h \mathbf{u}^{n+\frac{1}{2}}} \right\|_{L^2} \left\| \bar{e}_u^{n+\frac{1}{2}} \right\|_{L^6} \\ &\quad + \left\| \tilde{\mathbf{u}}_h^{n+\frac{1}{2}} \right\|_{L^\infty} \left\| \nabla \bar{e}_u^{n+\frac{1}{2}} \right\|_{L^2} \left\| \bar{\mathbf{u}}^{n+\frac{1}{2}} - \overline{\mathbf{R}_h \mathbf{u}^{n+\frac{1}{2}}} \right\|_{L^2} \\ &\leq C(\tau^2 + h^{r+1})^2 + \varepsilon \left\| \nabla \bar{e}_u^{n+\frac{1}{2}} \right\|_{L^2}^2 + C \left\| \bar{e}_u^{n+\frac{1}{2}} \right\|_{L^2}^2, \end{aligned}$$

where we have used (4.8), (4.17), (4.38),

$$\begin{aligned} \left\| \bar{\mathbf{u}}^{n+\frac{1}{2}} - \overline{\mathbf{R}_h \mathbf{u}^{n+\frac{1}{2}}} \right\|_{L^2} &= \left\| \frac{\mathbf{u}^{n+1} + \mathbf{u}^n}{2} - \frac{\widehat{\mathbf{R}_h \mathbf{u}^{n+1}} + \mathbf{R}_h \mathbf{u}^n}{2} \right\|_{L^2} \quad (\text{here use } \widehat{\mathbf{u}}^{n+1} = \mathbf{u}^{n+1}) \\ &\leq \left\| \frac{\mathbf{u}^{n+1} + \mathbf{u}^n}{2} - \frac{\mathbf{R}_h \mathbf{u}^{n+1} + \mathbf{R}_h \mathbf{u}^n}{2} \right\|_{L^2} + \left\| \frac{\mathbf{R}_h \mathbf{u}^{n+1} - \widehat{\mathbf{R}_h \mathbf{u}^{n+1}}}{2} \right\|_{L^2} \\ &\leq Ch^{r+1} + \frac{\tau}{4} \|\nabla_h(\mathbf{R}_h p^{n+1} - \mathbf{R}_h p^n)\|_{L^2} \quad (\text{by (4.27)}) \\ &\leq Ch^{r+1} + C\tau^2 \end{aligned}$$

and

$$\begin{aligned} \left\| \nabla \cdot \tilde{\mathbf{e}}_{\mathbf{u}}^{n+\frac{1}{2}} \right\|_{L^3} \left\| \bar{\mathbf{u}}^{n+\frac{1}{2}} - \overline{\mathbf{R}_h \mathbf{u}^{n+\frac{1}{2}}} \right\|_{L^2} &\leq Ch^{-1-\frac{d}{6}} \left\| \tilde{\mathbf{e}}_{\mathbf{u}}^{n+\frac{1}{2}} \right\|_{L^2} \left\| \bar{\mathbf{u}}^{n+\frac{1}{2}} - \overline{\mathbf{R}_h \mathbf{u}^{n+\frac{1}{2}}} \right\|_{L^2} \\ &\leq Ch^{-1-\frac{d}{6}} (h^{r+1} + \tau^2) \left\| \tilde{\mathbf{e}}_{\mathbf{u}}^{n+\frac{1}{2}} \right\|_{L^2} \quad (\text{by } \tau = \mathcal{O}(h)) \\ &\leq C \left\| \tilde{\mathbf{e}}_{\mathbf{u}}^{n+\frac{1}{2}} \right\|_{L^2}. \end{aligned}$$

The estimate for other terms of \mathcal{I}_4 is straightforward. Clearly, $\mathcal{I}_{4,1}$ and $\mathcal{I}_{4,5}$ are cancelled. By (4.9) and (4.38), we obtain

$$\mathcal{I}_{4,3} + \mathcal{I}_{4,4} + \sum_{k=6}^8 \mathcal{I}_{4,k} \leq C \left\| \tilde{\mathbf{e}}_{\mathbf{u}}^{n+\frac{1}{2}} \right\|_{L^2}^2 + \varepsilon \left\| \nabla \tilde{\mathbf{e}}_{\mathbf{u}}^{n+\frac{1}{2}} \right\|_{L^2}^2 + Ch^{2(r+1)} + C\tau^4.$$

Above all, we are led to

$$\mathcal{I}_4 \leq C \left\| \tilde{\mathbf{e}}_{\mathbf{u}}^{n+\frac{1}{2}} \right\|_{L^2}^2 + \varepsilon \left\| \nabla \tilde{\mathbf{e}}_{\mathbf{u}}^{n+\frac{1}{2}} \right\|_{L^2}^2 + Ch^{2(r+1)} + C\tau^4.$$

Similarly, we can rewrite \mathcal{I}_5 as

$$\begin{aligned} \mathcal{I}_5 &= -\mu \left\{ \left((\widetilde{\mathbf{H}}^{n+\frac{1}{2}} - \Pi_h \widetilde{\mathbf{H}}^{n+\frac{1}{2}}) \times (\nabla \times \widetilde{\mathbf{H}}^{n+\frac{1}{2}}), \bar{\mathbf{e}}_{\mathbf{u}}^{n+\frac{1}{2}} \right) + \left(\tilde{\mathbf{e}}_{\mathbf{H}}^{n+\frac{1}{2}} \times (\nabla \times \widetilde{\mathbf{H}}^{n+\frac{1}{2}}), \bar{\mathbf{e}}_{\mathbf{u}}^{n+\frac{1}{2}} \right) \right. \\ &\quad \left. + \left(\widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \times (\nabla \times (\widetilde{\mathbf{H}}^{n+\frac{1}{2}} - \Pi_h \widetilde{\mathbf{H}}^{n+\frac{1}{2}})), \bar{\mathbf{e}}_{\mathbf{u}}^{n+\frac{1}{2}} \right) + \left(\widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \times (\nabla \times \tilde{\mathbf{e}}_{\mathbf{H}}^{n+\frac{1}{2}}), \bar{\mathbf{e}}_{\mathbf{u}}^{n+\frac{1}{2}} \right) \right\} \\ &=: \sum_{k=1}^4 \mathcal{I}_{5,k}. \end{aligned}$$

By (4.12), we have

$$\begin{aligned} \mathcal{I}_{5,1} + \mathcal{I}_{5,2} &\leq \mu \left\| \widetilde{\mathbf{H}}^{n+\frac{1}{2}} - \Pi_h \widetilde{\mathbf{H}}^{n+\frac{1}{2}} \right\|_{L^2} \left\| \nabla \times \widetilde{\mathbf{H}}^{n+\frac{1}{2}} \right\|_{L^3} \left\| \bar{\mathbf{e}}_{\mathbf{u}}^{n+\frac{1}{2}} \right\|_{L^6} \\ &\quad + \mu \left\| \tilde{\mathbf{e}}_{\mathbf{H}}^{n+\frac{1}{2}} \right\|_{L^2} \left\| \nabla \times \widetilde{\mathbf{H}}^{n+\frac{1}{2}} \right\|_{L^3} \left\| \bar{\mathbf{e}}_{\mathbf{u}}^{n+\frac{1}{2}} \right\|_{L^6} \\ &\leq Ch^{2(r+1)} + \varepsilon \left\| \nabla \tilde{\mathbf{e}}_{\mathbf{u}}^{n+\frac{1}{2}} \right\|_{L^2}^2 + C \left\| \tilde{\mathbf{e}}_{\mathbf{H}}^{n+\frac{1}{2}} \right\|_{L^2}^2. \end{aligned}$$

With an application of integration by parts, $\mathcal{I}_{5,3}$ becomes

$$\mathcal{I}_{5,3} = \left(\bar{\mathbf{e}}_{\mathbf{u}}^{n+\frac{1}{2}} \times \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}}, \nabla \times (\widetilde{\mathbf{H}}^{n+\frac{1}{2}} - \Pi_h \widetilde{\mathbf{H}}^{n+\frac{1}{2}}) \right)$$

$$\begin{aligned}
&= \left(\nabla \times \left(\bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \times \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \right), \widetilde{\mathbf{H}}^{n+\frac{1}{2}} - \Pi_h \widetilde{\mathbf{H}}^{n+\frac{1}{2}} \right) \\
&\leq \left\| \nabla \times \left(\bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \times \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \right) \right\|_{L^2} \left\| \widetilde{\mathbf{H}}^{n+\frac{1}{2}} - \Pi_h \widetilde{\mathbf{H}}^{n+\frac{1}{2}} \right\|_{L^2} \\
&\leq \varepsilon \left\| \nabla \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right\|_{L^2}^2 + Ch^{2(r+1)},
\end{aligned}$$

where we have used (4.37) and

$$\begin{aligned}
&\left\| \nabla \times \left(\bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \times \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \right) \right\|_{L^2} \\
&= \left\| \left(\nabla \cdot \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \right) \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} - \left(\nabla \cdot \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right) \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} + \left(\widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \cdot \nabla \right) \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} - \left(\bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \cdot \nabla \right) \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \right\|_{L^2} \\
&\leq C \left\| \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \right\|_{W^{1,3}} \left\| \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right\|_{L^6} + C \left\| \nabla \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right\|_{L^2} \left\| \widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \right\|_{L^\infty}.
\end{aligned}$$

Therefore, the following bound is available for \mathcal{I}_5 :

$$\mathcal{I}_5 \leq Ch^{2(r+1)} + \varepsilon \left\| \nabla \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right\|_{L^2}^2 + C \left\| \bar{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right\|_{L^2}^2 - \mu \left(\widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \times \left(\nabla \times \bar{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right), \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right).$$

A bound for the truncation error term \mathcal{I}_6 is based on (4.9) and the regularity assumptions (3.5):

$$\mathcal{I}_6 \leq C(\tau^2 + h^{r+1})^2 + \varepsilon \left\| \nabla \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right\|_{L^2}^2.$$

With the above estimates, we obtain the following result from (4.42)

$$\begin{aligned}
&\frac{1}{2\tau} \left(\left\| \widehat{e}_{\mathbf{u}}^{n+1} \right\|_{L^2}^2 - \left\| e_{\mathbf{u}}^n \right\|_{L^2}^2 \right) + \frac{\nu}{2} \left\| \nabla \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right\|_{L^2}^2 - \left(e_p^n, \nabla \cdot \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right) \\
&\leq C \left\| \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right\|_{L^2}^2 + C(\ell_h \tau^2 + h^{r+1})^2 + C \left\| \bar{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right\|_{L^2}^2 - \mu \left(\widetilde{\mathbf{H}}_h^{n+\frac{1}{2}} \times \left(\nabla \times \bar{e}_{\mathbf{H}}^{n+\frac{1}{2}} \right), \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right). \tag{4.43}
\end{aligned}$$

Step 3: Estimate of the term $-\left(e_p^n, \nabla \cdot \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right)$ in (4.43). We rewrite (4.25) as

$$\frac{e_{\mathbf{u}}^{n+1} - \widehat{e}_{\mathbf{u}}^{n+1}}{\tau} = -\frac{1}{2} \nabla_h (e_p^{n+1} - e_p^n). \tag{4.44}$$

With the above equality and the fact that $\bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} = \frac{1}{2}(\widehat{e}_{\mathbf{u}}^{n+1} + e_{\mathbf{u}}^n)$, we have

$$\begin{aligned}
-\left(e_p^n, \nabla \cdot \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}} \right) &= -\frac{1}{2} \left(e_p^n, \nabla \cdot \widehat{e}_{\mathbf{u}}^{n+1} \right) \quad (\text{by (4.26)}) \\
&= \frac{1}{2} \left(\nabla_h e_p^n, \widehat{e}_{\mathbf{u}}^{n+1} \right) \\
&= \frac{\tau}{4} \left(\nabla_h e_p^n, \nabla_h (e_p^{n+1} - e_p^n) \right) \quad (\text{by (4.44)}) \\
&= \frac{\tau}{8} \left(\left\| \nabla_h e_p^{n+1} \right\|_{L^2}^2 - \left\| \nabla_h e_p^n \right\|_{L^2}^2 \right) - \frac{\tau}{8} \left\| \nabla_h (e_p^{n+1} - e_p^n) \right\|_{L^2}^2 \\
&= \frac{\tau}{8} \left(\left\| \nabla_h e_p^{n+1} \right\|_{L^2}^2 - \left\| \nabla_h e_p^n \right\|_{L^2}^2 \right) - \frac{1}{2\tau} \left\| e_{\mathbf{u}}^{n+1} - \widehat{e}_{\mathbf{u}}^{n+1} \right\|_{L^2}^2. \tag{4.45}
\end{aligned}$$

Step 4. By taking $\mathbf{l}_h = e_{\mathbf{u}}^{n+1}$ into (4.25), we arrive at

$$\frac{1}{2\tau} \left(\left\| e_{\mathbf{u}}^{n+1} \right\|_{L^2}^2 - \left\| \widehat{e}_{\mathbf{u}}^{n+1} \right\|_{L^2}^2 + \left\| e_{\mathbf{u}}^{n+1} - \widehat{e}_{\mathbf{u}}^{n+1} \right\|_{L^2}^2 \right) = 0, \tag{4.46}$$

in which (4.26) has been applied.

Step 5. A summation of (4.41), (4.43), (4.45), and (4.46) leads to

$$\begin{aligned} & \frac{\mu}{2\tau} \left(\|e_{\mathbf{H}}^{n+1}\|_{L^2}^2 - \|e_{\mathbf{H}}^n\|_{L^2}^2 \right) + \frac{\mu}{8\tau} \left(\|e_{\mathbf{H}}^{n+1} - e_{\mathbf{H}}^n\|_{L^2}^2 - \|e_{\mathbf{H}}^n - e_{\mathbf{H}}^{n-1}\|_{L^2}^2 \right) \\ & + \frac{1}{4\sigma} \|\nabla \times \tilde{e}_{\mathbf{H}}^{n+\frac{1}{2}}\|_{L^2}^2 + \frac{1}{4\sigma} \|\nabla \cdot \tilde{e}_{\mathbf{H}}^{n+\frac{1}{2}}\|_{L^2}^2 \\ & + \frac{1}{2\tau} \left(\|e_{\mathbf{u}}^{n+1}\|_{L^2}^2 - \|e_{\mathbf{u}}^n\|_{L^2}^2 \right) + \frac{\nu}{2} \|\nabla \bar{e}_{\mathbf{u}}^{n+\frac{1}{2}}\|_{L^2}^2 + \frac{\tau}{8} \left(\|\nabla_h e_p^{n+1}\|_{L^2}^2 - \|\nabla_h e_p^n\|_{L^2}^2 \right) \\ & \leq C \|\tilde{e}_{\mathbf{H}}^{n+\frac{1}{2}}\|_{L^2}^2 + C \|\tilde{e}_{\mathbf{H}}^{n+\frac{1}{2}}\|_{L^2}^2 + C \|\tilde{e}_{\mathbf{u}}^{n+\frac{1}{2}}\|_{L^2}^2 + C(\ell_h \tau^2 + h^{r+1})^2. \end{aligned} \tag{4.47}$$

An application of discrete Gronwall’s inequality indicates that there exists a positive constant τ_0 such that

$$\begin{aligned} & \|e_{\mathbf{H}}^{n+1}\|_{L^2}^2 + \tau \sum_{m=1}^n \|\nabla \times \tilde{e}_{\mathbf{H}}^{m+\frac{1}{2}}\|_{L^2}^2 + \tau \sum_{m=1}^n \|\nabla \cdot \tilde{e}_{\mathbf{H}}^{m+\frac{1}{2}}\|_{L^2}^2 \\ & + \|e_{\mathbf{u}}^{n+1}\|_{L^2}^2 + \tau \sum_{m=1}^n \|\nabla \bar{e}_{\mathbf{u}}^{m+\frac{1}{2}}\|_{L^2}^2 + \tau^2 \|\nabla_h e_p^{n+1}\|_{L^2}^2 \leq C(\ell_h \tau^2 + h^{r+1})^2, \end{aligned} \tag{4.48}$$

if $\tau < \tau_0$. By applying the Cauchy’s inequality

$$\|\nabla \times \tilde{e}_{\mathbf{H}}^{n+\frac{1}{2}}\|_{L^2}^2 \geq \frac{3}{8} \|\nabla \times e_{\mathbf{H}}^{n+1}\|_{L^2}^2 - \frac{1}{8} \|\nabla \times e_{\mathbf{H}}^{n-1}\|_{L^2}^2, \tag{4.49}$$

$$\|\nabla \cdot \tilde{e}_{\mathbf{H}}^{n+\frac{1}{2}}\|_{L^2}^2 \geq \frac{3}{8} \|\nabla \cdot e_{\mathbf{H}}^{n+1}\|_{L^2}^2 - \frac{1}{8} \|\nabla \cdot e_{\mathbf{H}}^{n-1}\|_{L^2}^2, \tag{4.50}$$

we further get the following result from (4.48)

$$\begin{aligned} & \|e_{\mathbf{H}}^{n+1}\|_{L^2}^2 + \tau \sum_{m=1}^n \|\nabla \times e_{\mathbf{H}}^{m+1}\|_{L^2}^2 + \tau \sum_{m=1}^n \|\nabla \cdot e_{\mathbf{H}}^{m+1}\|_{L^2}^2 \\ & + \|e_{\mathbf{u}}^{n+1}\|_{L^2}^2 + \tau \sum_{m=1}^n \|\nabla \bar{e}_{\mathbf{u}}^{m+\frac{1}{2}}\|_{L^2}^2 + \tau^2 \|\nabla_h e_p^{n+1}\|_{L^2}^2 \leq C(\ell_h \tau^2 + h^{r+1})^2. \end{aligned} \tag{4.51}$$

The above estimate implies that the induction assumption (4.36) could be recovered at $m = n + 1$ with $C_0^* \geq C$. Thus the mathematical induction is closed. By the projection estimates (4.9), (4.12), and (4.17), the error estimates (3.6) and (3.7) in Theorem 3.1 follow immediately.

□

Remark 4.5. In this work, we focus on the error estimates of the velocity and magnetic fields for a Crank–Nicolson finite element projection method. The error estimate of the pressure may be obtained by using the discrete inf-sup condition. From the numerical results in Section 4, we can see that the convergence for pressure is consistent with that for $\left(\tau \sum_{n=2}^N \|\nabla(\bar{\mathbf{u}}_h^{n-\frac{1}{2}} - \bar{\mathbf{u}}^{n-\frac{1}{2}})\|_{L^2}^2 \right)^{\frac{1}{2}}$.

5. NUMERICAL EXAMPLES

In this section, we present several numerical examples to illustrate our theoretical results in Theorems 3.1 and 3.6. For the sake of simplicity, numerical results are tested for two-dimensional problems in a unit square domain. All the numerical examples are computed by using FreeFEM++.

TABLE 1. Temporal convergence at $T = 1$.

τ	$e(\mathbf{u}_h)$	$e(\mathbf{H}_h)$	$e(\mathbf{p}_h)$	$e(\nabla \mathbf{u}_h)$	$e(\nabla \times \mathbf{H}_h)$
1/40	5.971×10^{-4}	1.862×10^{-3}	3.136×10^{-2}	1.167×10^{-2}	7.659×10^{-3}
1/80	1.495×10^{-4}	4.695×10^{-4}	8.487×10^{-3}	3.193×10^{-3}	1.906×10^{-3}
1/160	3.741×10^{-5}	1.179×10^{-4}	2.167×10^{-3}	8.176×10^{-4}	4.755×10^{-4}
Order	2.00	1.99	1.93	1.92	2.01

Example 5.1. First, we consider the MHD equations

$$\mu \partial_t \mathbf{H} + \sigma^{-1} \nabla \times (\nabla \times \mathbf{H}) - \mu \nabla \times (\mathbf{u} \times \mathbf{H}) = \mathbf{g}, \quad (5.1)$$

$$\partial_t \mathbf{u} + \mathbf{u} \cdot \nabla \mathbf{u} - \nu \Delta \mathbf{u} + \nabla p = \mathbf{f} - \mu \mathbf{H} \times (\nabla \times \mathbf{H}), \quad (5.2)$$

$$\nabla \cdot \mathbf{H} = 0, \quad \nabla \cdot \mathbf{u} = 0, \quad (5.3)$$

in $\Omega = [0, 1] \times [0, 1]$, with the initial and boundary conditions (1.4) and (1.5), where the source terms \mathbf{g} and \mathbf{f} are chosen correspondingly to the exact solutions

$$\begin{aligned} \mathbf{u} &= t^4 \begin{pmatrix} \sin^2(\pi x) \sin(2\pi y) \\ -\sin(2\pi x) \sin^2(\pi y) \end{pmatrix}, \\ \mathbf{H} &= t^4 \begin{pmatrix} -\sin(2\pi y) \cos(2\pi x) \\ \sin(2\pi x) \cos(2\pi y) \end{pmatrix}, \\ p &= t^4 \sin^2(2\pi x) \sin^2(2\pi y) - \frac{1}{4}. \end{aligned} \quad (5.4)$$

For simplicity, all the coefficients ν, σ, μ in (5.1)–(5.3) are chosen to be 1, and we take the final time as $T = 1$. Note that the above exact solutions \mathbf{u} and \mathbf{H} satisfy the divergence-free conditions and $\nabla p|_{\partial\Omega} = \mathbf{0}$.

We solve the MHD system (5.1)–(5.3) by the modified Crank–Nicolson finite element projection scheme (3.1)–(3.4) with a cubic finite element approximation for \mathbf{H} and \mathbf{u} , and a quadratic finite element approximation for p . Here, the system (3.1) and (3.2) is a coupled, but linearized system for \mathbf{H}_h^{n+1} and $\hat{\mathbf{u}}_h^{n+1}$. Thus we can solve the linear system directly by using a sparse solver in FreeFEM++. To investigate the convergence rate in time, we first choose $\tau = T/N$ with $N = 40, 80, 160$, with the spatial mesh size $h = 2\tau$. In Examples 5.1 and 5.2, we compute the following errors:

$$\begin{aligned} e(\mathbf{u}_h) &:= \|\mathbf{u}_h^N - \mathbf{u}^N\|_{L^2}, \quad e(\mathbf{H}_h) := \|\mathbf{H}_h^N - \mathbf{H}^N\|_{L^2}, \quad e(\mathbf{p}_h) := \|\mathbf{p}_h^N - \mathbf{p}^N\|_{L^2}, \\ e(\nabla \mathbf{u}_h) &:= \left(\tau \sum_{n=2}^N \|\nabla \hat{\mathbf{u}}_h^{n-\frac{1}{2}} - \nabla \hat{\mathbf{u}}^{n-\frac{1}{2}}\|_{L^2}^2 \right)^{\frac{1}{2}}, \quad e(\nabla \times \mathbf{H}_h) := \left(\tau \sum_{n=2}^N \|\nabla \mathbf{H}_h^n - \nabla \mathbf{H}^n\|_{L^2}^2 \right)^{\frac{1}{2}}. \end{aligned}$$

The numerical results at time $T = 1$ are presented in Table 1, which indicate that the proposed scheme has second-order convergence in time.

Then we solve the problem (5.1)–(5.3) by the modified Crank–Nicolson FEM scheme (3.1)–(3.4) with a sufficiently small temporal step size $\tau = 1/5000$, to focus on the spatial convergence rate. Again, a cubic finite element approximation for \mathbf{H} and \mathbf{u} is applied, combined with a quadratic finite element approximation for p . Here, we take $h = 1/10, 1/20, 1/40$. Numerical results at $T = 1$ are presented in Table 2. It is observed that the spatial convergences are of optimal orders, which are consistent with the theoretical analysis in Theorem 3.1.

TABLE 2. Spatial convergence at $T = 1$.

h	$e(\mathbf{u}_h)$	$e(\mathbf{H}_h)$	$e(\mathbf{p}_h)$	$e(\nabla \mathbf{u}_h)$	$e(\nabla \times \mathbf{H}_h)$
1/10	1.312×10^{-4}	2.481×10^{-4}	3.344×10^{-3}	2.386×10^{-3}	1.485×10^{-3}
1/20	8.184×10^{-6}	1.507×10^{-5}	5.186×10^{-4}	3.153×10^{-4}	1.727×10^{-4}
1/40	5.202×10^{-7}	9.320×10^{-7}	7.097×10^{-5}	4.068×10^{-5}	2.086×10^{-5}
Order	3.99	4.03	2.78	2.94	3.08

TABLE 3. Temporal convergence at $T = 1$.

τ	$e(\mathbf{u}_h)$	$e(\mathbf{H}_h)$	$e(\mathbf{p}_h)$	$e(\nabla \mathbf{u}_h)$	$e(\nabla \times \mathbf{H}_h)$
1/40	5.774×10^{-3}	1.912×10^{-3}	3.370×10^{-2}	1.210×10^{-2}	7.766×10^{-3}
1/80	1.411×10^{-4}	4.822×10^{-4}	1.085×10^{-2}	3.949×10^{-3}	1.932×10^{-3}
1/160	3.500×10^{-5}	1.211×10^{-4}	3.563×10^{-3}	1.301×10^{-3}	4.819×10^{-4}
Order	2.02	1.99	1.62	1.61	2.01

TABLE 4. Spatial convergence at $T = 1$.

h	$e(\mathbf{u}_h)$	$e(\mathbf{H}_h)$	$e(\mathbf{p}_h)$	$e(\nabla \mathbf{u}_h)$	$e(\nabla \times \mathbf{H}_h)$
1/10	1.304×10^{-4}	2.481×10^{-4}	1.996×10^{-3}	2.306×10^{-3}	1.485×10^{-3}
1/20	7.878×10^{-6}	1.507×10^{-5}	2.504×10^{-4}	2.859×10^{-4}	1.727×10^{-4}
1/40	4.867×10^{-7}	9.321×10^{-7}	3.112×10^{-5}	3.560×10^{-5}	2.086×10^{-5}
Order	4.02	4.03	3.00	3.01	3.08

Example 5.2. Second, we solve the MHD model (5.1)–(5.3) by the scheme (3.1)–(3.4) with the source terms \mathbf{g} and \mathbf{f} chosen correspondingly to the following exact solutions

$$\begin{aligned}
 \mathbf{u} &= t^4 \begin{pmatrix} \sin^2(\pi x) \sin(2\pi y) \\ -\sin(2\pi x) \sin^2(\pi y) \end{pmatrix}, \\
 \mathbf{H} &= t^4 \begin{pmatrix} -\sin(2\pi y) \cos(2\pi x) \\ \sin(2\pi x) \cos(2\pi y) \end{pmatrix}, \\
 p &= t^4 \sin(2\pi x) \sin(2\pi y),
 \end{aligned} \tag{5.5}$$

where we can see that $\nabla p|_{\partial\Omega} \neq \mathbf{0}$. Here, we choose the same time step sizes and mesh sizes as those used in Example 5.1, and compute the errors and convergence. The numerical results are given in Tables 3 and 4, which are consistent with the theoretical results in Theorem 3.1.

Example 5.3. Third, we test the energy stability of the proposed scheme by solving the problem (1.1)–(1.5) in $\Omega = [0, 1] \times [0, 1]$ with $\mathbf{J} = \mathbf{f} = \mathbf{0}$ (\mathbf{J} denotes a scalar function in \mathbb{R}^2) and $T = 1$. Here, all the coefficients

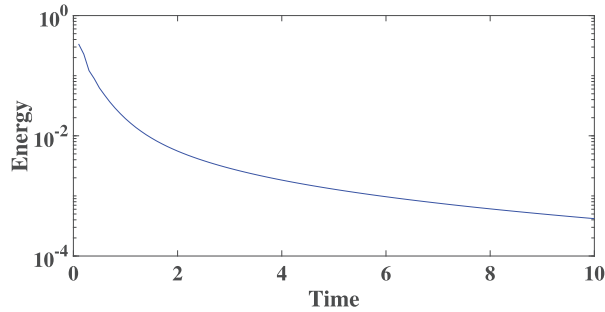


FIGURE 1. Energy of the MHD system at each time level.

ν, σ, μ in (1.1)–(1.5) are chosen to be 1 and the initial values are chosen as:

$$\begin{aligned} \mathbf{u}_0 &= \begin{pmatrix} \sin^2(\pi x) \sin(2\pi y) \\ -\sin(2\pi x) \sin^2(\pi y) \end{pmatrix}, \\ \mathbf{H}_0 &= \begin{pmatrix} -\sin(2\pi y) \cos(2\pi x) \\ \sin(2\pi x) \cos(2\pi y) \end{pmatrix}, \\ p_0 &= \sin(2\pi x) \sin(2\pi y). \end{aligned} \quad (5.6)$$

We solve the problem by the proposed scheme (3.1)–(3.4) with a quadratic finite element approximation for \mathbf{H} and \mathbf{u} , combined with a linear finite element approximation for p . The time step size and spatial mesh size are chosen as $\tau = 1/10$ and $h = 1/50$, respectively. We define the energy function as $E_h^n := \|\mathbf{u}_h^n\|_{L^2}^2 + \|\mathbf{H}_h^n\|_{L^2}^2 + \frac{1}{4}\|\mathbf{H}_h^n - \mathbf{H}_h^{n-1}\|_{L^2}^2 + \frac{\tau^2}{4}\|\nabla_h p_h^n\|_{L^2}^2$. The energy evolution curve, up to the final time $T = 10$, is displayed in Figure 1, which clearly indicates the energy dissipation property, consistent with the theoretical result in Theorem 3.6.

6. CONCLUSION

In this paper, we propose a fully discrete decoupled finite element projection method for the incompressible magnetohydrodynamic equations (1.1)–(1.3). The primary difficulties are associated with the nonlinear and coupled nature of the problem. In this work, a modified Crank–Nicolson method is used for the temporal discretization, and appropriate semi-implicit treatments are adopted for the approximation of the fluid convection term and two coupled terms. Then a linear system with variable coefficients is presented and its unique solvability is theoretically proved by the fact that the corresponding homogeneous equations only admit zero solutions. One prominent advantage of the scheme is associated with a decoupling approach in the Stokes solver, which computes an intermediate velocity field based on the pressure gradient at the previous time step, and then enforces the incompressibility constraint *via* the Helmholtz decomposition of the intermediate velocity field. As a result, this decoupling approach greatly reduces the computation of the MHD system. Furthermore, the energy stability analysis and error estimates in the discrete $L^\infty(0, T; L^2)$ norm are provided for the scheme, in which the decoupled Stokes solver needs to be carefully estimated. Several numerical examples are presented to demonstrate the robustness and accuracy of the proposed scheme. The extension of the energy stable projection methods and its error estimates to two-phase MHD models will be investigated in the future.

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